# EXP-T

## a relativistic multireference coupled cluster program

v 1.8.0

User manual

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# 1 Introduction

The EXP-T program package is designed for high-precision modeling of molecular electronic structure using the relativistic Fock space coupled cluster method (FS-RCC). Features and program components are listed in Sec. 1.1 and 1.2.

The authors will be grateful for any comments or suggestions: exp-t-program@googlegroups.com https://groups.google.com/d/forum/exp-t-program

## 1.1 Features of EXP-T

The EXP-T package does not include subroutines for solving the (Dirac-) Hartree-Fock equations and subsequent four-index transformation, so molecular integrals have to be imported from third party electronic structure packages. Currently EXP-T is interfaced to the DIRAC program package [1, 2], thus getting access to the wide variety of Hamiltonians and property operators implemented therein.

Electronic structure models available in EXP-T:

- single-point energy calculations with any point groups and (nearly) all Hamiltonians, implemented in DIRAC (4c-DC, X2Cmmf, 2c-ECP, non-relativistic);
- ground state energy calculations (the 0h0p sector): CCSD, CCSD(T), CCSDT-n (n=1,2,3), CCSDT models;
- the FS-CCSD method for excited states is implemented for the 0*h*1*p*, 1*h*0*p*, 1*h*1*p*, 0*h*2*p*, 2*h*0*p*, 0*h*3*p*, 1*h*2*p* Fock space sectors;
- FS-CC models for excited states accounting for triples (CCSDT-n (n=1,2,3), CCSDT) are implemented for the 0h1p, 1h0p, 1h1p, 0h2p, 2h0p, 0h3p Fock space sectors.

The summary of all electronic structure models implemented in the EXP-T program system is given in the table:

FS sector	CCSD	$\mathrm{CCSD} + \mathrm{T}(3)$	CCSDT-1	CCSDT-2	CCSDT-3	CCSDT
0h0p	+	+	+	+	+	+
0h1p	+	t	t	t	t	t
1h0p	+	t	t	t	t	t
1h1p	+	t	t	t	t	t
0h2p	+	t	t	t	t	t
2h0p	+	—	t	t	t	t
0h3p	+	t	t	t	t	t
1h2p	t	_	_	—	—	—

"+" - implemented, "-" - not implemented, "t" - currently being tested

Features for property calculations and analysis of wavefunctions:

- analytic density matrices and property calculations (including  $\mathcal{T}$ -odd operators) for the ground-state CCSD and CCSD(T) models;
- finite-field [3] and direct [4] approaches to transition moments calculations;
- quasidiabatization of SO-coupled states and extraction of spin-orbit matrix elements [5].
- "quasi-natural" spinors (NS) and natural transition spinors (NTS) for non-trivial Fock space sectors;

• one-electron properties via the interface to the OneProp program by L. V. Skripnikov [6].

Features to deal with the intruder state problem:

- "dynamic" energy denominators shifts as a solution of the intruder-state problem [5];
- Padé extrapolation to the zero-shift limit [7];
- intermediate Hamiltonian technique for incomplete main model spaces (IH-IMMS) [8];

Features for high-performance calculations:

- OpenMP parallelization (for shared-memory systems);
- parallel calculations on NVIDIA GPUs using the CUDA platform [9].

#### **1.2** Program components

expt.x Input processing + all CC calculations.

heffman.x Manipulations with effective Hamiltonian matrices: Padé extrapolations, finite-field transition moments calculations, quasidiabatization of SO-coupled states.

expt\_transform\_natorb.x

Transformation of "quasi-natural" spinors to the AO basis.

expt\_diatomic.x

Rotational-vibrational levels and properties of diatomic molecules.

expt2pam.x Tool for convenient basis set and pseudopotential input.

expt2spectrum.py

Tool for convenient extraction of calculated information on transitions and spectra from EXP-T output files.

#### 1.3 Citation

We kindly ask you to acknowledge any use of the EXP-T program system that results in published material using the following citation:

A. V. Oleynichenko, A. Zaitsevskii, E. Eliav. Towards high performance relativistic electronic structure modelling: the EXP-T program package. *Commun. Comput. Inf. Sci.* 1331, 375 (2020). doi: 10.1007/978-3-030-64616-5\_33

In addition, credit must be given to the corresponding papers which describe the implementation and the underlying methodological developments:

- Triples in the low Fock space sectors: [10]
- The 0h3p Fock space sector: [11]
- Finite-field transition matrix elements (general): [3]

- Finite-field transition matrix elements (the  $0h0p \rightarrow 1h1p$  excitations): [12]
- Spin-orbit coupling calculations: [5]
- Dynamic denominator shifts (real): [5]
- Dynamic denominator shifts (imaginary, or their real simulation): [13]
- Padé extrapolation of effective Hamiltonian series: [7]
- Intermediate Hamiltonian for incomplete main model spaces (IH-IMMS): [8]
- Direct method of transition matrix elements calculations: [4]

## 1.4 Credits

EXP-T is based on ideas and design solutions of the suite of FS-CC programs written by Ephraim Eliav and Uzi Kaldor.

# 2 Installation from source code

## 2.1 Compiling EXP-T

EXP-T is currently oriented at Unix-like operating systems.

Tools required for compiling the EXP-T package from source code:

• C and Fortran compilers. Currently supported compiler systems are

gcc/gfortran GNU compilers icc/ifort Intel compilers

 cmake (version 3.0.2 or higher) and make utilities: https://cmake.org/ https://www.gnu.org/software/make/

Optional dependencies:

- Implementation of the BLAS/LAPACK linear algebra libraries (OpenBLAS [14] or Intel MKL [15] are recommended);
- NVIDIA drivers and CUDA Toolkit are required to perform parallel calculations on GPU: https://www.nvidia.com/Download/index.aspx?lang=en-us https://developer.nvidia.com/cuda-toolkit

Download the EXP-T source code as a \*.tar.xz file and unpack it:

```
tar xvf expt-X.Y.Z.tar.xz
```

Go to the EXP-T home directory, create the build directory and make it the current working dir:

mkdir build cd build

Compilation (Intel compilers):

```
CC=icc FC=ifort cmake .. make
```

Compilation (GNU compilers):

```
CC=gcc FC=gfortran cmake .. make
```

We recommend specifying compilers explicitly (since the default C compiler may be something other than gcc).

The CMake utility will try to locate BLAS/LAPACK libraries on your computer and link them to the project. If there are no pre-installed libraries on your machine, the internal OpenBLAS will be compiled and linked to EXP-T.

When using Intel MKL together with GNU compilers, you will probably need to explicitly specify the path to the MKL home directory by setting the environment variable MKLROOT. For example:

#### export MKLROOT=/opt/intel/mkl

In case of successful compilation, the executable file expt.x will appear in the build directory. We strongly recommend that compilation is followed by testing (see Sect 2.4). Do not forget to add the directory containing the binaries to the PATH environment variable!

#### 2.1.1 Compiling on Mac OS X

The default C compiler for Mac OS X is clang; this compiler was not tested, so we strongly recommend to use gcc instead:

CC=gcc FC=gfortran cmake .. make

We strongly recommend not to use GNU compilers installed with Homebrew. Instead, download the GNU compilers suitable for your version of Mac OS X from:

http://hpc.sourceforge.net/

In order to use a BLAS/LAPACK implementation other than the one attached to the EXP-T package, you can use the following manual:

https://pheiter.wordpress.com/2012/09/04/howto-installing-lapack-and-blas-on-mac-os/

However, the Netlib implementation described there can perform several times slower than the Open-BLAS implementation included in EXP-T.

## 2.2 Compiling with CUDA

Compilation of EXP-T does not require the user to have a CUDA-compatible GPU device. If the CUDA toolkit is missing, the CUDA code will simply be excluded from the compilation process.

EXP-T is adapted for parallel calculations on NVIDIA GPUs (the CUDA technology [9]). Notes:

- 1. At the moment, the OpenMP+CUDA hybrid model as well as multi-GPU support is not implemented. Only one GPU can be used for calculations.
- 2. non-professional GPUs for gaming are not suitable for CUDA calculations (except for Kepler cards).
- 3. CUDA code in EXP-T was tested only for CUDA 9.1 and 10.0.
- 4. Only Intel and GNU compilers were tested for compatibility with CUDA.

To compile EXP-T with CUDA:

- 1. Make sure your graphics card supports CUDA. CUDA-compatible GPUs are listed here: https://developer.nvidia.com/cuda-gpus
- Install proprietary NVIDIA drivers (not the nouveau driver!). https://www.nvidia.com/Download/index.aspx?lang=en-us. For Linux you can find drivers in your distribution's repository.
- Install CUDA toolkit https://developer.nvidia.com/cuda-downloads or from distribution's repository.
- 4. Make sure that all drivers and CUDA toolkit are installed and work correctly. Compile the test suite in the samples directory of the CUDA home directory and run some of them. The most important are the tests employing the CUBLAS linear algebra library.
- 5. Compile EXP-T as described in Sec. 2.1. All CUDA tools will be located automatically by CMake.
- 6. Make sure that the results produced by EXP-T do not depend on the use of CUDA.

### 2.3 Troubleshooting

This section describes the most common problems which can occur during the compilation process. In case of any questions or problems, please don't hesitate to contact us. Don't forget to attach the output file (in case of problems with calculations) or the file with the output of the make VERBOSE=1 command (in case of problems with compilation).

Google Groups: https://groups.google.com/d/forum/exp-t-program

E-mail:

exp-t-program@googlegroups.com

#### 2.3.1 Outdated version of CMake

We highly recommend using CMake version 3.0.2 or higher, otherwise, problems may occur when building the internal OpenBLAS. The following command will show what version of CMake is installed:

cmake --version

If CMake version is lower than 3.0.2, download the binary distribution from their official website https://cmake. org/download/. After installation do not forget to add the directory containing the CMake binary files to the PATH environment variable.

Installation of CMake binary files does not require root privileges.

## 2.4 Testing

Test suite is located in the test directory. The testing system is written in the Python programming language (must be pre-installed on your machine). Either Python 2 or Python 3 can be used.

In order to run the testing suite, run the following command:

make test

## 3 Running EXP-T

See Sec. 6 for a description of an additional utility programs included in the EXP-T package.

A typical coupled cluster calculation consists of two sequential steps:

1. SCF and integral transformation within DIRAC (Sect. 3.1);

2. CC calculation within EXP-T (Sect. 3.2).

#### 3.1 Precomputing spinors and molecular integrals: the DIRAC step

The interface to DIRAC was tested only for the DIRAC17, DIRAC18 and DIRAC19 releases. DIRAC stores transformed molecular integrals in the following binary files:

MRCONEE information about SCF calculation, symmetry, occupation numbers, one-electron integrals;

MDCINT symmetry-unique nonzero two-electron integrals;

MDPROP one-electron property integrals.

Run DIRAC and save the files containing the transformed integrals to the working directory using the --get option:

#### pam --noarch --mol=<mol-file> --inp=<inp-file> --get="MRCONEE MDCINT MDPROP"

If DIRAC starts in parallel mode (MPI), it is necessary to choose the "scheme 4" transformation algorithm. By default, DIRAC uses the "scheme 6" algorithm which produces multiple MDCINT files [16] and currently EXP-T requires that all integrals are stored in a single file. In order to accomplish this, add the following lines to the DIRAC input file:

\*\*MOLTRA .SCHEME 4

To calculate properties, transition moments and intensities without resorting to the finite-field scheme property integrals transformed to the basis of molecular spinors are also required. We recommend explicitly asking DIRAC to calculate and transform these integrals by adding the .PRPTRA keyword together with the list of the required property operators to DIRAC input files:

If the molecular-mean-field X2C Hamiltonian (X2Cmmf) was used at the DIRAC stage, one **must** specify the **x2cmmf** flag in the EXP-T input file! (see also page 23).

## 3.2 Coupled cluster calculation: the EXP-T step

Run an EXP-T job using the following command:

expt.x [options] <input-file>

The EXP-T output (including error messages) is flushed to the UNIX standard output (stdout). Use the output redirection operator ">" in case of long time calculations:

```
nohup expt.x <input-file> > <output-file> &
```

Examples of DIRAC and EXP-T input files are given in Sec. 5.

## 3.3 Command-line arguments

Use the expt.x --help command to print a list of available command-line arguments:

Usage: expt.x [OPTION...] <input-file>
expt -- A Fock-Space Multireference Relativistic Coupled-Cluster Program

-n,	no-clean	Do not clean scratch directory on exit
		(use this option to keep cluster amplitudes etc)
-s,	scratch=PATH	Path to scratch directory (default: ./scratch)
-?,	help	Print this help list and exit
	usage	Print a short usage message and exit
-V,	version	Print program version and exit

Mandatory or optional arguments to long options are also mandatory or optional for any corresponding short options.

Please report bugs to <exp-t-program@googlegroups.com>.

## 4 Input files syntax and keywords

See Sec. 6 for the description of additional utility programs included in the EXP-T package.

The input file format may change in future versions.

The EXP-T input file contains a list of instructions. Each instruction is placed on a separate line and consists of a keyword and a list of arguments. Single-line comments start with '#' and continue until the end of the line. Input is case insensitive. Examples of EXP-T input files are given in Sec. 5. Molecular geometry, basis set etc have to be described only once in the DIRAC input files at the molecular integrals calculation stage (not in EXP-T input files).

The detailed description of the keywords used in the EXP-T input file is provided below. For convenience, keywords are grouped into several sections. The order of instructions in the input file can be arbitrary.

#### General settings

optional comment. title Syntax: title <quoted-string> Default: no title Example: title "my first ccsd calculation" allows one to force EXP-T to use the complex arithmetic everywhere. arith Syntax: arith complex print level. print Syntax: print ( low || medium || high || debug || <quoted-string> ) where <quoted-string> can be: "eff config" effective configurations of spinors for each electronic state; expansions of model vectors in terms of model determinants; "model vectors" information about model space determinants. "model space" One is allowed to combine several **print** instructions. Default: print medium Example: print low print "eff config" degen\_thresh energy levels are considered as degenerate if the energy gap between them is lower than degen\_thresh. Syntax: degen\_thresh <real thresh>

> Units: atomic Default: 1e-8

spinor\_labels manual assignment of symbolic names (labels) of specific spinors (orbitals). Symbolic name is printed together with spinor symmetry and energy. This option can be useful when studying composition of electronic states.

Syntax:

```
spinor_labels
   . . .
   <integer spinor-number> <string spinor-label>
   . . .
end
```

Default: disabled.

Example: labels of atomic spinors:

```
spinor_labels
    2 2s_1/2
    3 2p_1/2
    4 2p_3/2
    5 2p_3/2
end
```

#### Model

sector specifies the target Fock space sector. To the moment the 0h0p, 0h1p, 1h0p, 1h1p, 0h2p, 2h0p, 0h3p, 1h2p FS sectors are implemented (h – holes, p – particles).

Syntax: sector <integer H>h<integer P>p

Default: sector OhOp (vacuum)

model CC model (approximation to the cluster operator).

Default: model ccsd

Notes:

- Keywords ccsd(t) and ccsd+t(3) are equivalent.
- The codes implementing the CCSD+T(3), CCSDT-n and CCSDT models in the nontrivial sectors and all CC models in the 0h3p sector are experimental (currently being tested), to be used with care.
- The current implementation of the CCSD(T) model for the 0h0p sector is not very computationally efficient.
- The CCS and CCD models are implemented simply by setting all  $T_2$  (CCS) or  $T_1$  (CCD) amplitudes to zero, this does not make computation faster, and exist primarily for the purposes of testing and debugging.

occ spinor occupation numbers. This keyword can be used to setup calculations with a "high-spin" vacuum state [17]. This keyword is alternative to the occ\_irreps keyword (see below).

Syntax: occ <list of the 0 and 1 digits separated by spaces>

Default: occupation numbers are read from MO integrals files

Example: 5 electrons and 10 spinors: occ 1 1 0 0 0 1 1 1 0 0

CC calculations with high-spin vacuum states were tested for the 0h0p FS sector only.

It is rather convenient to use the dirac\_mointegral\_export routine [18] from the DIRAC package to generate string with occupation numbers.

occ\_irreps overall number of electrons in each irrep. Overall irrep occupations are ignored if occupation
numbers of individual spinors are given (by the occ keyword, see above). Names of irreps are
listed in Appendix C.

Syntax: occ\_irreps <pairs irrep:occ>

Example: occupation numbers for the ground triplet state of the  $O_2$  molecule:

# for gerade irreps: occ\_irreps [1/2g+]:3 [1/2g-]:2 [3/2g+]:0 [3/2g-]:1 # for ungerade irreps: occ\_irreps [1/2u+]:2 [1/2u-]:2 [3/2u+]:1 [3/2u-]:1

Default: occupation numbers are read from MO integrals files.

#### Active space

- nacth number of highest occupied spinors (active holes)
- nactp number of lowest virtual spinors (active particles)

It is convenient to setup the active space by the nacth/nactp keywords when exploring electronic states of molecules in a range of geometries (potential energy surfaces, transition moment functions etc). Using the nacth/nactp keywords one can specify either overall number of active spinors (holes or particles) or number of active spinors in each irreducible representation.

Syntax: nacth <integer dim>
nactp <integer dim>
or
nacth <pairs irrep:number-of-spinors>
nactp <pairs irrep:number-of-spinors>

Default: nacth 0, nactp 0 (no active spinors).

Example: consider the low-lying electronic states of the La atom. Those are dominated by the  $6s^{n}5d^{m}$  configurations, generating even electronic states. So it seems reasonable to include the  $6s_{1/2}$ ,  $6d_{3/2}$  and  $6d_{5/2}$  atomic spinors of the La<sup>3+</sup> cation into the active space and achieve target electronic states in the 0h3p Fock space sector. In order to make the input file easy to

read, one can use two sequential **nactp** directives, one for the "negative" projections of  $m_j$ , and one for the "positive" ones:

nactp [1/2g+]:3 [3/2g+]:2 [5/2g+]:1nactp [1/2g-]:3 [3/2g-]:2 [5/2g-]:1One can extend the scope of achievable electronic states by inclusion of the 6p and 4f spinors: nactp [1/2u+]:4 [3/2u+]:3 [5/2u+]:2 [7/2u+]:1nactp [1/2u-]:4 [3/2u-]:3 [5/2u-]:2 [7/2u-]:1

active active (i. e. valence) space specification by energy range. Active spinors must have one-electron energies lying in the range  $\varepsilon_i \in [\varepsilon_{min}; \varepsilon_{max}]$  (regardless of their symmetry). Setting the active space using the active energy instruction is convenient only for atomic calculations.

Syntax: active energy <real eps\_min> <real eps\_max>

Default: no active spinors

Example: active energy -1.0 5.0

#### Intermediate Hamiltonian for incomplete main model spaces (IH-IMMS)

Intermediate Hamiltonian technique for incomplete main model spaces (IH-IMMS) can be used to achieve convergence of iterative solution of CC equations. The first paper [8] on the IH-IMMS method concerns only the special case of atomic systems. Structure of incomplete model spaces for molecules is simpler than for the special case of atoms; recent molecular applications with detailed discussions on IH-IMMS for molecules can be found in [4] (the ThO molecule) and [19] (the AcF molecule).

The basic idea of the method is to split the whole model space into the "main" and "intermediate" subspaces. Slater determinants  $|\Phi_M\rangle$  which span the main model space must be chosen to describe the bulk of wavefunctions of target electronic states. FS-CC equations for amplitudes corresponding to excitations from main determinants remain unchanged. In contrast, equations for amplitudes from intermediate ("buffer") determinants  $|\Phi_I\rangle$  are solved with shifted energy denominators to suppress the intruder state problem. For each cluster amplitude corresponding to the excitation  $I \to K$  from  $|\Phi_I\rangle$  a shift parameter  $S_{IK}$  is calculated as:

$$S_{IK} = \overline{E}_0 - E_I,$$

where  $E_I$  stands for the energy of the  $|\Phi_I\rangle$  model-space determinant and  $\overline{E}_0$  stands for the some "frontier" energy. This value is used to calculate a shift for energy denominator corresponding to a cluster amplitude using one of the formulas (4)–(7). All model space determinants  $|\Phi_M\rangle$  with energy  $E_M \leq E_0$  are considered as main ones. Determinants  $|\Phi_I\rangle$  with energy  $E_I > E_0$  are intermediate.

By default, IH-IMMS is disabled. IH settings are specified using the complex instruction ih\_imms:

ih\_imms

```
<option-1>
<option-2>
<option-3>
...
end
```

To set up the IH-IMMS calculation:

1. split the set of active spinors into several subspaces. In the current version of EXP-T only spinor energies can be used to select these subspaces;

- 2. define model Slater determinants which will be considered as the "main" ones. Sets of "main" determinants are defined via the numbers of particles (holes) in each spinor subspace;
- 3. (optional) define the "frontier" energy  $\overline{E}_0$ . By default, frontier energy is calculated automatically.

The use of IH-IMMS for the simple case of the carbon atom is considered in details in Sect. 5.4. The following keywords can be used inside the *ih\_imms* block:

sectors list of Fock space sectors in which the IH-IMMS technique will be applied. This keyword is mandatory.

Syntax: sectors <list-of-sector-labels>

Default: no target FS sectors.

Example: sectors 0h1p 0h2p

subspace defines subspace of active spinors with energies  $\varepsilon \in [\varepsilon_{\min}, \varepsilon_{\max}]$ . Subspaces are used to define composition of main model space determinants via the main\_occ keyword. One should use several subspace options in the input file to define more than one subspace of active spinors. This keyword is mandatory, at least one subspace must be specified.

Syntax: subspace energy <real emin> <real emax>

Default: no subspaces are defined.

Example: subspace energy -1.0 -0.5

main\_occ defines subspace of main model space determinants. The argument of this option contains the number of particles (holes) in the 1st, 2nd, ... subspaces of active spinors. One should use several main\_occ options in the input file to define more than one subspace of main model space determinants. This keyword is mandatory, at least one subspace of determinants must be specified.

Syntax: main\_occ <list-of-integers>

Default: no main model spaces are defined.

Example: imagine that we have to define model space determinants with 2 electrons on active spinors belonging to the 1st spinor subspace, 1 electron on spinors from the 2nd subspace and 0 electrons on the other active spinors (such a model determinant belongs to the 0h3p sector). The proper option is as follows:

main\_occ 2 1 0

Syntax: shift\_type ( none || real || realimag || imag || taylor )

Default: shift\_type realimag

**npower** attenuation parameter used in the shifting formula (see Eq. (4)-(7)).

Syntax: npower <int n>

Default: npower 5

scale scaling factor for the shift parameter S:

 $S \rightarrow S' = \alpha S$ 

Syntax: scale <real factor> Default: scale 1.0 (no scaling)

frontier\_energy "frontier" energy  $\overline{E}_0$  used to calculate shift for the given (intermediate) determinant.

Syntax: frontier\_energy <real energy>

Default: frontier energy is evaluated automatically as the maximum energy of main model space determinants:

 $\overline{E}_0 = \max_{\Phi_M \in \text{main}} E_M, \qquad E_M = \langle \Phi_M | H_0 | \Phi_M \rangle.$ 

#### **Properties**

analyt\_prop analytic calculation of coupled cluster expectation values of arbitrary property operators (including  $\mathcal{T}$ -odd ones). Matrix elements of property operators must be stored in the DIRAC's MDPROP unformatted file. A-equations, analytic density matrix and properties are currently available only for the 0h0p sector (vacuum state) and the CCSD and CCSD(T) models.

Syntax: analyt\_prop <property-name>

Default: disabled.

Example: calculation of x,y,z-components of the dipole moment.

analyt\_prop XDIPLEN
analyt\_prop YDIPLEN
analyt\_prop ZDIPLEN

density calculation of analytic density matrix in the 0h0p sector. Prints occupation numbers of CC natural orbitals. Analytic density matrices are currently implemented for the CCSD and CCSD(T) models.

Syntax: density 0h0p

- Default: disabled.
- hermit perform symmetric orthogonalization of model vectors. Orthogonalized model vectors are used only to calculate properties and approximate density matrices; matrices of effective Hamiltonians flushed to HEFF files (see Sec. 6.1) remain non-Hermitian.

Syntax: hermit

Default: "hermitization" is disabled.

mdprop enable calculations of transition property matrix elements (not only dipole moments). Property operator matrix elements in the basis of molecular spinors are to be loaded from the DIRAC's MDPROP unformatted file. To avoid huge outputs, use options nroots or roots\_cutoff to restrict the set of target roots (ot the expt\_spectrum.py utility to conveniently extract data for the specific transitions, see Sect. 6.4). Note that the currently available maximum truncation level is quadratic, which might be not enough for the systems with strong dynamic correlations (and thus large cluster amplitudes).

The details on the direct method of transition matrix element calculations are given in [4]. Within this approach property matrix element  $O_{if}$  for the pair of electronic states  $\psi_i$  and  $\psi_f$  can be calculated via the relation [20] (scheme non-hermitian):

$$O_{if} = \langle \tilde{\psi}_i^{\perp \perp} | (\Omega^{\dagger} \Omega)_{closed} \Omega^{\dagger} O \Omega | \tilde{\psi}_f \rangle \cdot N_i N_f^{-1}, \tag{1}$$

where  $\tilde{\psi}_i^{\perp\perp}$  stands for the left model vector and the normalization factors are defined as

$$N_i = \left\langle \psi_i | \psi_i \right\rangle^{1/2} = \left\langle \tilde{\psi}_i | \Omega^{\dagger} \Omega | \tilde{\psi}_i \right\rangle^{1/2} \tag{2}$$

(the same for  $N_f$ ). This definition of the effective property operator leads to the non-Hermitian property matrix,  $O_{if} \neq O_{fi}^*$ . Any hermitization procedure can be applied.

There is also an alternative definition (scheme hermitian):

$$O_{if} = \langle \tilde{\psi}_i | \Omega^{\dagger} O \Omega | \tilde{\psi}_f \rangle \cdot N_i^{-1} N_f^{-1}$$
(3)

which is inherently Hermitian [20]. Expressions  $\Omega^{\dagger}\Omega$  and  $\Omega^{\dagger}O\Omega$  are evaluated using the FS-CC Ansatz  $\Omega = \{e^T\}$  and have to be artificially truncated at some power of T (the approx keyword). It is natural to use the same level of truncation for both these expressions. Finally, the transformation properties with respect to the molecular point group operations of the specific property operator should be explicitly given (the sym keyword).

Syntax: mdprop "<property-name>" [sym <irrep>] [approx <n> <m>] [scheme hermitian | non-hermitian]

	1	C 1	1	1	1	• 1 . •
C TTM	cummotru	of the	nronorty	onorator 1	indor	concideration
БVШ	SVIIIIIGUIV	OI UNC	DIODELUV	0 $0$ $0$ $1$ $0$ $0$ $1$ $0$ $1$ $0$ $0$ $1$ $0$ $0$ $1$ $0$ $0$ $1$ $0$ $0$ $0$ $0$ $0$ $0$ $0$ $0$ $0$ $0$	maci	consideration.
j j			r r v	- <u>r</u>		)

- approx truncation level for the numerator  $\Omega^{\dagger}O\Omega$  and the metric operator  $\Omega^{\dagger}\Omega$ ; 0 for full neglect of terms containing cluster amplitudes, 1 for linear approximation, 2 for quadratic approximation;
- scheme particular form of the effective operator to be used, non-hermitian for (1) or hermitian for (3).

Default: disabled.

Example: calculation of the whole set of possible dipole transition moments for a linear diatomic heteronuclear molecule (the  $C_{\infty v}$  point group). Terms which are quadratic in cluster amplitudes in both numerator and denominator are included (approx 2 2). The hermitian formula is used. The list of instructions will be as follows:

```
mdprop "XDIPLEN" sym 1u+ approx 2 2 scheme hermitian
mdprop "XDIPLEN" sym 1u- approx 2 2 scheme hermitian
mdprop "YDIPLEN" sym 1u+ approx 2 2 scheme hermitian
mdprop "YDIPLEN" sym 1u- approx 2 2 scheme hermitian
mdprop "ZDIPLEN" sym 0u approx 2 2 scheme hermitian
```

Matrix elements of some operators calculated by DIRAC can be purely imaginary even if the molecular point group allows one to use real arithmetics. For this reason we recommend to thoroughly check the diagnostics in the EXP-T output. If matrix elements of such an operator are to be calculated, we recommend to solve amplitude equations in real arithmetic, then switch on the complex one using the arith complex instruction and calculate property matrix elements of interest.

The code for properties is currently experimental; please, we kindly ask you to report all failures and obviously incorrect results.

mstdm enable calculations of the model-space approximations to transition dipole moments (wavefunctions are represented by model vectors) [21, 22, 12]). To avoid huge outputs, use options nroots or roots\_cutoff to restrict the set of target roots.

Note that left model vectors are used in bra:

 $d_{if,\eta} = \langle \tilde{\psi}_i^{\perp \perp} | d_\eta | \tilde{\psi}_f \rangle \qquad \eta = x, y, z$ 

and if the "hermitization" of the effective Hamiltonian matrix is disabled, the TDM matrix will be (normally slightly) non-Hermitian  $(|d_{if}| \neq |d_{fi}|)$ .

This approach typically overestimates transition moments by 30-50%, hence we recommend using it only for semiquantitative estimations, e. g. detection of the most intensive transitions and discerning between  $0^+$  and  $0^-$  states (such E1-transitions are strictly forbidden).

Syntax: mstdm

Default: disabled.

txtprop enable calculations of arbitrary properties. Property operator matrix elements is loaded from two formatted files. To avoid huge outputs, use options nroots and roots\_cutoff to restrict the set of target roots. Other options are the same as for the mdprop keyword.

Syntax: txtprop <string file\_re> <string file\_im> [sym <irrep>] [approx <n> <m>]
[scheme hermitian | non-hermitian]

Default: disabled.

Example:

mdprop PropIntsRe.txt PropIntsIm.txt

The code for properties is currently experimental; please, we kindly ask you to report all failures and obviously incorrect results.

natorb calculation of "quasi-natural" spinors (NS) and transition spinors (NTS).

Note: density matrices  $d_{pq}^{IJ} = \langle \tilde{\psi}_I^{\perp \perp} | a_p^{\dagger} a_q | \tilde{\psi}_J \rangle$  are constructed with model vectors instead of true wavefunctions.

```
Syntax: natorb [<rep>:<state> | <rep1:state1>-<rep2:state2>] ...
```

Electronic states for which NS and NTS should be calculated are specified by the <rep>:<state> pair, where:

<re> name of irreducible representation in square braces. See Sect. C for the list of irrep names used in EXP-T.

<state> sequential number of the state in the irrep (1, 2, ...).

Pair of electronic states separated by a hyphen shows that NTS between these states are to be calculated. You can ask EXP-T to calculate NS and NTS for arbitrary number of states (or pairs of states).

Default: calculation of NS and NTS is disabled.

Examples:

- NS for the ground and the first excited states of the RaCl molecule [23]: natorb [1/2+]:1 [1/2+]:2
- NTS for the (1)0<sup>+</sup> (2)0<sup>+</sup> transition in the RbCs molecule:
   natorb [0]:1-[0]:3

(note that in the latter example the [0]:2 state is excluded, since it is the  $(1)0^{-}$  state).

#### Iterative solution and convergence

**conv** specifies the requested convergence threshold (in amplitudes):

 $\left||t_{K}^{(n+1)}| - |t_{K}^{(n)}|\right| < \texttt{conv} \quad \forall K \quad \rightarrow \text{converged}$ 

Syntax: conv <real thresh>

Default: 1e-9.

damping enable damping in the given FS sector. Amplitudes obtained at the *n*-th step will be mixed with amplitudes from the (n-1)-th step using the damping factor  $\alpha$ :

$$t_K^{(n)'} = \alpha t_K^{(n-1)} + (1-\alpha) t_K^{(n)}$$

Damping can aid the convergence of equations in some non-pathological cases (when the lack of convergence is caused by amplitude oscillations rather than intruder states). Note that in other cases damping can significantly slow down convergence.

Syntax: damping <H>h<P>p <integer last\_step> <real factor>

<H>h<P>p FS sector in which these damping parameters will be applied

last\_step at which iteration damping will be disabled

factor damping factor  $\alpha$ . Can be negative.

Default: disabled in all sectors

Example: damping 0h2p 100 0.5

diis enables the DIIS extrapolation technique [24, 25].

Syntax: diis [ ( off || triples || <integer dim> ) ]

Disable DIIS: diis off

Enable DIIS for triples: diis triples

Default: diis 10 (DIIS is enabled for singles and doubles, disabled for triples)

It is highly recommended to enable DIIS for triples if the CCSDT-n or CCSDT models are employed.

div\_thresh specifies the divergence threshold (in amplitudes). Solution of amplitude equation is diverged if for some  $t_K$  holds:

 $\left||t_{K}^{(n+1)}|-|t_{K}^{(n)}|\right|>\texttt{div\_thresh}\quad\rightarrow\text{diverged}$ 

May be useful if few initial steps of the iterative solution are divergent.

Syntax: div\_thresh <real thresh> Default: 1e+3.

**flush** indicates that cluster amplitudes must be flushed to disk every N iterations. Can be useful for very time-consuming calculations with poor convergence. Allows one to change the convergence acceleration strategy.

Syntax: flush <integer N> iter Default: no flush. Example: flush 5 iter **nroots** specifies the number of roots (electronic states) of interest in each irrep (in the target Fock space sector). Model vectors analysis will be performed only for the **nroots** lowest states.

Syntax: nroots <pairs irrep:number-of-roots>

Default: all roots are of interest.

Example: consider the electronic states of the RbCs molecule, corresponding to the first three dissociation limits (see, for example, [5, 26] and references therein). In the Hund's case a these electronic states are  $X^{1}\Sigma^{+}$ ,  $a^{3}\Sigma^{+}$ ,  $b^{3}\Pi$ ,  $A^{1}\Sigma^{+}$ ,  $c^{3}\Sigma^{+}$ ,  $B^{1}\Pi$ . In the Hund's case c these states are classified according the projection of the total electronic angular momentum and can be rearranged as follows:  $(1-3)0^{+}$ ,  $(1-3)0^{-}$ , (1-4)1, (1)2 states. The names of irreps used in EXP-T are given in Appendix C (the point group of RbCs is  $\overline{C_{\infty v}}$ ). All output and analysis of model vectors will be restricted to these 11 states using the following command: nroots [0]:6 [1+]:4 [1-]:4 [2+]:1 [2-]:1 (note that as in DIRAC  $0^{+}$  and  $0^{-}$  states belong to the same irrep).

maxiter maximum number of iterations allowed to solve the CC equations.

Syntax: maxiter <integer max> Default: maxiter 50.

reuse indicates that sorted integrals and cluster amplitudes from the previous calculation should be reused. Cluster amplitudes will be used as an initial guess. If EXP-T fails to locate the files containing amplitudes, they will be recalculated.

Syntax: reuse <list-of-arguments>

Possible values of arguments:

integrals	load one- and two-electron integrals (from a scratch directory);
1-integrals	one-electron integrals only;
2-integrals	two-electron integrals only;
amplitudes	load amplitudes for all FS sectors occuring in the calculation;
OhOp, Oh1p,	1h0p, 1h1p, 0h2p, 2h0p, 0h3p, 1h2p
	load amplitudes for the given sector only

Default: no reuse

Note: files with converged amplitudes are flushed to the scratch directory after each calculation (see also Appendix B). To prevent the deletion of files with amplitudes or integrals it is necessary to use the --no-clean command-line option (see Sect. 3.3).

Example: reuse integrals 0h0p 0h1p

roots\_cutoff specifies the upper energy bound for the electronic states to be analyzed and printed. Model vectors analysis and properties calculations will be performed only for the states with energy lower than roots\_cutoff. Allowed units of energy are eV, cm<sup>-1</sup> or Hartrees.

Syntax: roots\_cutoff <real max\_energy> ( ev || cm || au )

Default: all roots are of interest.

Example: roots\_cutoff 30000 cm

shift sets the denominator shift parameters in the  $n_h h n_p p$  Fock space sector. shift <H>h<P>p <int n> <real shift\_T1> <real shift\_T2> [<real shift\_T3>] attenuation parameter (see eqs. (4) and (7));

- **shift\_T1** shift value for the  $T_1^{n_h h \ n_p p}$  cluster operator (singles). Must be omitted in the 2h0p and 0h2p FS sectors. It is recommended to set  $\texttt{shift_T1} = 0.0$  in the 1h1p sector, unless you really understand the consequences of using nonzero values. Note that  $\texttt{shift_T1}$  value in this sector does not affect the resulting energies;
- shift\_T2 shift value for  $T_2^{n_h h \ n_p p}$  (doubles);
- shift\_T3 shift value for  $T_3^{n_h h n_p p}$  (triples). Must be used only for CC models including non-perturbative triples.

Default: n = 3, all shifts = 0.0 Examples: shift 0h1p 3 -0.3 -0.6 -0.9 shift 0h1p 3 -0.6 shift 1h1p 4 0.0 -0.5

n

shifttype indicates which formula for energy denominator shifts should be used [5, 7, 27].

shifttype is the same for all FS sectors. If shifttype is not equal to none, you must specify the shift parameters for at least one FS sector (see keyword shift; except for 0h0p).

Syntax: shifttype ( none || real || realimag || imag || taylor ) Possible values:

none	no shift
real	real shift

$$D'_{K} = D_{K} + S \left(\frac{S}{D_{K} + S}\right)^{n} \tag{4}$$

realimag real simulation of imaginary shift [13]

$$D'_{K} = D_{K} + \frac{S^{2}}{D_{K}} \left(\frac{S^{2}}{D_{K}^{2} + S^{2}}\right)^{n}$$
(5)

imag imaginary shift [13]

$$D'_{K} = D_{K} + iS \left(\frac{|S|}{|D_{K} + iS|}\right)^{n} \tag{6}$$

taylor extrapolated intermediate Hamiltonian-like shift [28]

$$D'_{K} = (D_{K} + S) \frac{\left(1 - \frac{S}{D_{K} + S}\right)}{1 - \left(\frac{S}{D_{K} + S}\right)^{n+1}}$$
(7)

#### Default: shifttype none

Shift is never applied in the 0h0p sector!

skip indicates which sectors will be skipped in the FS-CC calculation. Can be useful in very costly calculations (like CCSDT) in order to avoid recalculation of the first step in already converged sectors. To be used with the **reuse** option.

Syntax: skip <list-of-sector-symbols>

Default: no skip.

Example: skip 0h0p 0h1p

## Interfaces

integrals sets paths to files containing transformed molecular integrals.

```
Syntax: integrals <string "1-el int-s"> <string "2-el int-s">
    [<string "properties int-s">]
```

Arguments:

- one-electron integrals file;
- two-electron integrals file;
- properties transformed integrals file.

Default: integrals MRCONEE MDCINT MDPROP

Example: integrals ../MRCONEE-Cinfv ../MDCINT-Cinfv /home/user/MDPROP

oneprop imports one-electron property integrals from the MDPROP unformatted file. The one-electron operator  $\hat{O}_1$  is added to the Fock operator at the integral sorting stage:

 $\hat{f}' = \hat{f} + \lambda \hat{O}_1$ 

where  $\lambda$  is some perturbation parameter. EXP-T does not assume Kramers symmetry, so both  $\mathcal{T}$ -odd and  $\mathcal{T}$ -even operators are allowed.

```
Syntax: oneprop <real L_re> <real L_im> <quoted-string property-name>
```

L\_re, L\_im real and imaginary parts of  $\lambda$ ;

property-name string identifier used to encode property matrix in the MDPROP file (can be also found in the DIRAC output file).

Default: disabled.

Example: add interaction with an external electric field:

oneprop 1e-5 0.0 "ZDIPLEN"

x2cmmf since DIRAC's unformatted files do not contain any information about the relativistic Hamiltonian used, the x2cmmf flag must be used if the molecular-mean-field X2C Hamiltonian (X2Cmmf) was employed at the DIRAC stage.

Syntax: x2cmmf

## Memory management

compress enables compression of all data written to disk. The LZ4 algorithm [29] is employed (it provides the fastest decompression to date). Compression can slow down calculations slightly (depends on the particular harddrive), but the disk space used can be reduced dramatically (up to 3 times).

Syntax: compress

Default: disabled (except for disk\_usage=4)

disk\_usage indicates which data should be stored on the disk. Note that very intensive disk usage will slow down calculations; however, disk usage is hardly avoidable in case of large tasks.

Syntax: disk\_usage <integer mode>

Possible values of <mode>:

- 0 all data is stored in RAM;
- 1 tensors of rank  $\geq 6$  are stored on disk ( $T_3$  amplitudes etc);
- 2 tensors of rank  $\geq 6$  and  $\langle pp | | pp \rangle$  (4 particles) two-electron integrals are stored on disk;
- 3 tensors of rank  $\geq 6$ ,  $\langle pp | | pp \rangle$  and  $\langle *p | | pp \rangle$  (3 particles) two-electron integrals are stored on disk;
- 4 the same as 3, + compression of all data. Is suitable for extremely large tasks only.

Default: 2 ( $\langle pp | | pp \rangle$  and tensors of rank  $\geq 6$  are stored on disk).

**memory** specifies the amount of memory that EXP-T can use for the job (just for *dynamically* allocated memory).

Syntax: memory <real size> ( mb || gb ) Min value: 10 Mb Default: memory 1024 mb (1 Gb)

tilesize the whole set of one-electron functions (spinors or spin-orbitals) is divided into blocks: (1) by
irrep; (2) then into sub-blocks with size not exceeding tilesize (actually the same as "tiles"
in NWChem-TCE):



The N-dimensional arrays of integrals and amplitudes (tensors) also turn out to be divided into blocks containing  $\sim (\texttt{tilesize})^N$  elements:



tilesize must be large enough to place at least one N-dimensional array of size  $(tilesize)^N$  in RAM. We recommend using as large tilesize as possible, otherwise memory management overheads can become enormous.

Syntax: tilesize <integer size>

Default: tilesize 100

## Parallel execution

**nthreads** number of OpenMP threads. Note that scaling with respect to number of threads is much better for large tasks (> 400 spinors).

Syntax: nthreads <integer n\_omp\_threads>

Default: nthreads 1 (sequential execution).

At the moment parallelization is not used at the integral sorting stage, hence OpenMP can accelerate only the amplitude equations solution stage.

openmp the same as nthreads.

cudaenables parallel calculations on NVIDIA GPU (only a single GPU is supported at the moment).<br/>Requires nthreads 1 (the OpenMP+CUDA hybrid model is not implemented yet).Syntax: cudaDefault: CUDA disabled.

## 5 Running typical calculations

The input and output files discussed in this section can be found in the *examples* directory of the EXP-T distro.

### 5.1 Ground state energy with CCSD: the CO molecule

For input/output files see examples/CO\_ccsd.

Consider the calculation of the CCSD correlation energy of the CO molecule in the cc-pVDZ [30] basis set.

```
DIRAC input files:
   (CO.mol)
DIRAC
CO molecule, R = 2.132 bohrs
symmetry C1
С
    2
          0
        6.
               1
С
          0.0
                              0.0
                    0.0
LARGE BASIS cc-pVDZ
        8.
               1
         0.0
                              2.132
0
                   0.0
LARGE BASIS cc-pVDZ
FINISH
   (TRA.inp)
**DIRAC
.TITLE
CO molecule
.WAVE FUNCTION
.4INDEX
**WAVE FUNCTION
.SCF
*SCF
. CLOSED
14
. ERGCNV
1E-12
**MOLTRA
. ACTIVE
all
*END OF INPUT
```

We run DIRAC with the following command:

#### pam --inp=TRA --mol=CO --get="MRCONEE MDCINT" --noarch

Note that the --get option is used so that pam exports the transformed molecular integrals (stored in the MRCONEE and MDCINT files) from DIRAC's scratch directory to the working directory. The -noarch option is used in order to prevent creation of a tgz archive.

In some cases, it can be convenient to rename the files containing the integrals (this is not mandatory):

```
mv MRCONEE MRCONEE-C1
mv MDCINT MDCINT-C1
   Now the EXP-T input file input-C1 is set up:
# CO molecule, CCSD/cc-pVDZ
# nonrelativistic, symmetry C1
# task title (of type "string")
title "CO/CCSD/cc-pVDZ"
# print level, default medium
print medium
# max number of iterations
maxiter 30
# convergence threshold (by cluster amplitudes)
conv 1e-9
# target Fock space sector
sector OhOp
# CC model: ccsd, ccsdt-1, etc
model ccsd
# import integrals from
integrals MRCONEE-C1 MDCINT-C1
```

Running EXP-T:

expt.x input-C1

The string with the correlation energy can be found at the end of the output. Note that by default EXP-T flushes all output simply to stdout.

SCF reference	energy =	-112.820480227130
CCSD correlation	energy =	-0.298117912040
Total CCSD	energy =	-113.118598139170

### 5.2 Simple atomic FS-CCSD calculation: the Pb atom

For input/output files see examples/Pb\_fs\_ccsd.

We now turn to simple atomic relativistic calculations employing the FS-CCSD approach. Since wave functions of most low-lying electronic states of one-electron atoms and atomic ions are effectively singlereference, it is not very interesting to apply the multireference methodology to such cases. In contrast, electronic states of the "two-electron" atoms can hardly be simulated within any other approaches except those based on the multireference wave functions. Consider for example the  $6s^26p^2$  low-lying states of the Pb atom. The reference state will be Pb<sup>2+</sup>  $6s^2$ , and one should select 6p-spinors as the active ones (overall 6 spinors). Pb atom is a purely relativistic system, and one should use four-component Hamiltonian or relativistic pseudopotential approach. The latter allows one to incorporate bulk of the Breit interaction and QED corrections. Here the semilocal valence version of the N.S. Mosyagin's RECP for Pb is used (see [31] and references therein).

Input file for the FS-CCSD calculation will be as follows:

sector 0h2p nactp 6 maxiter 500 memory 110 gb conv 1e-7 disk\_usage 0 nthreads 8 model ccsd

After the FS-CCSD calculation is finished, one can compare the calculated FS-CCSD excitation energies of Pb with their experimentally measured counterparts (7819, 10650, 21458 and 29467 cm<sup>-1</sup>, respectively [32]):

@       1       -0.8258538771       .       0.00000000       0.000000       1       0g         @       2       -0.7916687782       .       0.9302239316       7502.761992       3       0g       1g+         @       3       -0.7784232250       .       1.2906537953       10409.824894       5       0g       1g+       1g-	
@ 3 -0.7784232250 1.2906537953 10409.824894 5 0g 1g+ 1g- 2g	
	;+ 2g-
@ 4 -0.7298818976 2.6115306047 21063.414836 5 0g 1g+ 1g- 2g	+ 2g-
© 5 -0.6892150215 3.7181326781 29988.762480 1 0g	

The remaining error can be significantly reduced by inclusion of triple excitations [10].

## 5.3 Electronic states of the HgH<sup>2+</sup> molecular ion. Energy denominator shifts technique

For input/output files see examples/HgH2+\_shifts.

In almost every problem involving the exploration of potential energy surfaces using the FS-CC method in a fairly wide range of internuclear distances, the intruder state problem arises. The presence of intruder states usually manifests itself as (but does not not simply arise from!) the appearance of small energy denominators in amplitude equations; the presence of small denominators makes the iterative procedure unstable. A straightforward solution of the intruder-state problem can be obtained by modifying (shifting) all energy denominators in the problematic Fock space sector(s) in such a way that the shifts are significant for ill-defined (nearly zero or positive) energy denominators and remain negligibly small for large negative denominators [5, 7]. Note, however, that this does not imply any modification of the right-hand side of the amplitude equations and, thus, introduces an additional (although controllable) approximation.

The denominator shift technique is implemented in EXP-T; different formulas are available (see Sect. 4, keyword shifttype). The example discussed below employs the simplest *real* shifts.

Consider the HgH<sup>2+</sup> molecular ion at R = 2.0 bohr (we partially follow the paper of N. S. Mosyagin et al [33]). The FS-CC equations cannot be solved at all without using denominator shifts or intermediate Hamiltonian technique. Let the active space be comprised of the 8 lowest virtual spinors. The EXP-T input file will look like this:

```
# (1) HgH2+ ion, FSCC scheme: HgH3+ -> HgH2+
# (2) hamiltionian: 2-comp gatchina ECP
# (3) symmetry Cinfv
# (4) sector 0h1p
title "Test HgH3+ -> HgH2+/gatchina ECP/Cinfv"
maxiter 200
conv 1e-9
active -10 -0.5
sector 0h1p
# import integrals from:
integrals MRCONEE-Cinfv MDCINT-Cinfv
# shifts of denominators:
# real shift:
shifttype real
# shift parameters:
# 1. <n> -- attenuation parameter (here: 3)
# 3. <S1> -- shift for all S1 amplitudes (here: -0.5)
# 3. <S2> -- shift for all S2 amplitudes (here: -1.0)
shift 0h1p 3 -0.5 -1.0
```

Approaches to the selection of shift parameters are discussed in [5, 7]. You can also perform IH-IMMS calculations or do Pade extrapolation of the series of effective Hamiltonian matrices to the zero shift limit in order to minimize distortions introduced by the shift [7] (see also Sect. 6.1).

### 5.4 Intermediate Hamiltonian for incomplete main model spaces: the C atom

For input/output files see examples/C\_atom\_ih-imms.

Let us consider low-lying electronic states of the C atom. Using the NIST ASD database https://www.nist.gov/pml/atomic-spectra-database we find that for the states below 80000 cm<sup>-1</sup> the following configurations dominate in wave functions:

- $2s^22p^2;$
- $2s^12p^3;$
- $2s^22p^13s^1;$
- $2s^22p^13p^1;$
- $2s^22p^13d^1;$
- $2s^22p^14s^1$ .

Within the Fock space coupled cluster formalism, electronic states dominated by the  $2s^12p^3$  configuration (e. g. the  ${}^5S_2^o$  state, 33735 cm<sup>-1</sup>) can be accessed only in 1h3p or 0h4p sectors. Implementation of corresponding CC models is not available to date, thus we will not consider these states in the further discussion. In contrast, other states of the carbon atom can be easily obtained in the 0h2p Fock space sector provided that the  $2s^2$  state of the C<sup>2+</sup> ion is considered as the Fermi vacuum. For the sake of simplicity we will not also consider electronic states involving shells with principal quantum number n = 4 (the generalization of the IH-IMMS procedure to such states is straightforward).

At this point we have several useful observations:

- active space should comprise at least 2p, 3s, 3p, 3d spinors of the C<sup>2+</sup> ion;
- only the 2p shell can be doubly occupied in the target electronic states;
- all target electronic states involve configurations with at least one 2*p*-electron;
- we can expect that electronic states dominated by the  $3s^2$ ,  $3s^13p^1$ , etc configurations will be very high in energy (and even lie well above the IP = 90820 cm<sup>-1</sup>). These states arise inevitably in the 0h2p sector since the "classical" FS-CC is formulated for quasicomplete model spaces. Moreover, these *intruder states* will destroy convergence of amplitude equations;
- the linear span of these "malicious" determinants can be considered as the *intermediate space*;
- main model space should include only determinants belonging to the  $2p^2$ ,  $2p^13s^1$ ,  $2p^13p^1$ ,  $2p^13d^1$  configurations. These determinants will be dominating in the model vectors representing target electronic states of carbon.

We start actual calculation with the molecular integral transformation in DIRAC: (C.mol)

```
C atom
doubly augmented triple-zeta basis set (d-aug-cc-pVTZ)
symmetry Dinfh
С
    1
        6.
              1
     0.000000000
С
                      0.000000000
                                       0.000000000
LARGE EXPLICIT
                  4
                        2
                             2
                                  1
                                        1
 s functions
#
F 10 2
       8236.0000000
                         0.00053100
                                       -0.00011300
       1235.00000000
                         0.00410800
                                       -0.00087800
        280.8000000
                         0.02108700
                                       -0.00454000
         79.27000000
                         0.08185300
                                       -0.01813300
         25.59000000
                         0.23481700
                                       -0.05576000
          8.99700000
                         0.43440100
                                       -0.12689500
          3.31900000
                         0.34612900
                                       -0.17035200
          0.90590000
                         0.03937800
                                        0.14038200
          0.36430000
                        -0.00898300
                                        0.59868400
          0.12850000
                         0.00238500
                                        0.39538900
F
   4
      0
          0.90590000
          0.12850000
          0.04402000
          0.01510000
 p functions
  5
F
     1
         18.71000000
                         0.01403100
          4.13300000
                         0.08686600
```

```
1.2000000
                         0.29021600
          0.38270000
                         0.50100800
          0.12090000
                         0.34340600
  4 0
F
          0.38270000
          0.12090000
          0.03569000
          0.01050000
# d functions
F
   4 0
          1.09700000
          0.31800000
          0.1000000
          0.03140000
# f functions
F
  3 0
          0.76100000
          0.26800000
          0.09440000
FINISH
   (TRA.inp)
**DIRAC
.TITLE
Carbon atom
.4INDEX
.WAVE FUNCTION
**HAMILTONIAN
.X2Cmmf
.GAUNT
**WAVE FUNCTION
.SCF
*SCF
. CLOSED
4 0
. ERGCNV
1E-12
**MOLTRA
. SCHEME
 4
.ACTIVE
all
*END OF INPUT
```

Note that the .X2Cmmf keyword is used, thus it also must be used in the EXP-T input file.

Command line:

pam --inp=TRA --mol=C --noarch --get="MRCONEE MDCINT MDPROP"

Let us now look at the DIRAC output. We find that one-electron energies of our active spinors are (in a.u.):

 $\varepsilon_{2p} \sim -0.86$ 

 $\varepsilon_{3s} \sim -0.36$  $\varepsilon_{3p} \sim -0.28$  $\varepsilon_{3d} \sim -0.23$ 

We split these spinors into two subsets. Spinors from the first subset can be doubly occupied in the main model space determinants, and spinors from the second subset can be only singly occupied. In some cases it can be useful to define the third subset of spinors which must be unoccupied in all main determinants. This procedure resembles that used to define active space in the RASSCF method. For the C atom these subsets are:

1. 2p;

2. 3s, 3p, 3d.

In terms of these two subspaces of active spinors the main model space determinants can be defined using the following directives inside the *ih\_imms* block in the EXP-T input file:

- main\_occ 2 0: determinants in which both electrons occupy spinors from the first subspace, the  $2p^2$  configuration;
- main\_occ 1 1: one electron occupy spinor from the first subspace (2p), the other one occupy spinor from the second subspace (3s, 3p or 3d). These determinants correspond to the  $2p^13s^1$ ,  $2p^13p^1$  and  $2p^13d^1$  configurations.

We will rely on the automatic determination of the **frontier\_energy** parameter (by default, the uppermost energy of main model space determinants is used). Finally, the EXP-T input file will be as follows:

```
# relativistic CC model
sector 0h2p
model ccsd
nactp 24
x2cmmf
# intermediate Hamiltonian parameters.
# spinor subspaces are defined via energy ranges (in a.u.)
ih imms
  sectors 0h2p
  subspace energy -1.0 -0.5 # 2p spinors
  subspace energy -0.5 -0.2 # 3s, 3p, 3d spinors
                             # 2p^2 determinants
  main_occ 2 0
                             # 2p3s, 2p3p, 2p3d determinants
  main_occ 1 1
end
# memory
memory 4 gb
disk_usage 0
# convergence
maxiter 300
diis 300
conv 1e-9
```

#### # print compositions of model vectors in terms of # effective occupation numbers of active spinors print "eff config"

Let us take a look at the table with calculated energy levels in the 0h2p sector in the EXP-T output file:

Level Re(eigenvalue) ... Rel eigv, cm-1 % main deg symmetry

		0	0		0	0		0					
0	1	-1.2955562077	0.000000	100.0	1	0g							
0	2	-1.2954943219	13.582372	100.0	3	0g	1g+	1g-					
0	3	-1.2953712703	40.589069	100.0	5	0g	1g+	1g-	2g+	2g-			
0	4	-1.2503086362	9930.694080	99.9	5	0g	1g+	1g-	2g+	2g-			
0	5	-1.1923569864	22649.611052	99.8	1	0g							
0	6	-1.0208724411	60286.118415	99.9	1	0u							
0	7	-1.0207910577	60303.980008	99.9	3	0u	1u+	1u-					
0	8	-1.0206248742	60340.453071	99.9	5	0u	1u+	1u-	2u+	2u-			
0	9	-1.0089659280	62899.295993	99.9	3	0u	1u+	1u-					
0	10	-0.9902128248	67015.126416	100.0	3	0g	1g+	1g-					
0	11	-0.9824896577	68710.165670	100.0	3	0g	1g+	1g-					
0	12	-0.9824017917	68729.450018	100.0	5	0g	1g+	1g-	2g+	2g-			
0	13	-0.9822669729	68759.039338	100.0	7	0g	1g+	1g-	2g+	2g-	3g+	3g-	
0	14	-0.9751409822	70323.013514	100.0	3	0g	1g+	1g-					
0	15	-0.9723223910	70941.622775	100.0	1	0g							
0	16	-0.9722734205	70952.370551	100.0	3	0g	1g+	1g-					
0	17	-0.9721863073	70971.489684	100.0	5	0g	1g+	1g-	2g+	2g-			
0	18	-0.9616074654	73293.277111	99.8	5	0g	1g+	1g-	2g+	2g-			
0	19	-0.9497229655	75901.623363	99.9	1	0g							
0	20	-0.9322863070	79728.527553	100.0	5	0u	1u+	1u-	2u+	2u-			
0	21	-0.9307044048	80075.714943	100.0	5	0u	1u+	1u-	2u+	2u-			
0	22	-0.9306318565	80091.637473	100.0	7	0u	1u+	1u-	2u+	2u-	3u+	3u-	
0	23	-0.9305012392	80120.304651	100.0	9	0u	1u+	1u-	2u+	2u-	3u+	3u-	
0	24	-0.9303011549	80164.218067	100.0	3	0u	1u+	1u-					
0	25	-0.9302612163	80172.983578	100.0	5	0u	1u+	1u-	2u+	2u-			
0	26	-0.9302131715	80183.528205	100.0	7	0u	1u+	1u-	2u+	2u-	3u+	3u-	
0	27	-0.9264080659	81018.652356	100.0	7	0u	1u+	1u-	2u+	2u-	3u+	3u-	
0	28	-0.9244722663	81443.511260	100.0	5	0u	1u+	1u-	2u+	2u-			
0	29	-0.9243950805	81460.451580	100.0	3	0u	1u+	1u-					
0	30	-0.9243547369	81469.305985	100.0	1	0u							
0	31	-0.9218925156	82009.701090	99.9	3	0u	1u+	1u-					
0	32	-0.5545617414	162629.487352	9.6	1	0g							
0	33	-0.5088229529	172667.991098	0.0	1	0u							
0	34	-0.5088105565	172670.711776	0.0	3	0u	1u+	1u-					
0	35	-0.5087857581	172676.154402	0.0	5	0u	1u+	1u-	2u+	2u-			
י ר	 Ionizet	tion notential Obon	-> 0h1n =	= 10 981	V و 9	=	221	575	16 cr	n^_1			
1	[onizat	tion potential Oh1p	-> 0h0p =	= 24.272	0 eV	=	195	766.	56 cr	n^_1			
_										_			

The column named "% main" contains total weight of main model space determinants in model vectors describing the given electronic states. It can be seen that these weights are very close to 100% for all target (or main) electronic states. Intermediate states (with weighs of main model space determinants not exceeding several percents) are also obtained as a by-product. However, they can hardly be considered as physical ones (compare their energies with the  $IP_1 \sim 90000 \text{ cm}^{-1}$ ). The greater the weight of main

model space determinants, the less the dependence on the parameters of the applied denominator shifting scheme.

It is worth noting that in some cases convergence of amplitude equations can be achieved only by the introduction of additional active spinors, which are unoccupied in all main model space determinants. These spinors serves as buffer ones. In calculations accounting for triples such an extension of active space can result in an increase of computational cost. So it is recommended to thoroughly choose the upper bound of the active space to make it as small as possible. However, this is not a problem for calculations employing the CCSD model only. Introduction of additional buffer states also greatly improves stability of energy levels with respect to the shifting parameters (see Figures 3 and 4 in [8] for the case of electronic states of the  $Lu^+$  ion).

#### 5.5 Finite-field transition dipole moments calculations: the Rb atom

For input/output files see examples/Rb\_atom\_TDMs\_finite\_field.

See [22, 34, 3] for detailed discussions on the finite field (FF) technique for transition dipole moments. The algorithm for any FF calculation is as follows:

- solve the HF problem for the non-perturbed Hamiltonian (no external fields);
- transform the molecular integrals to the basis of "non-perturbed" one-electron functions;
- perform two FS-CC calculations with these integrals at external fields strengths -F and +F;
- estimate TDMs are estimated using the finite-difference-type formula:

$$\boldsymbol{d}_{\eta,if}^{FF} = (E_i - E_f) \lim_{\substack{F_\eta \to 0 \\ F_{\eta'} = 0, \eta' \neq \eta}} \frac{\langle \psi_i^{\perp \perp}(-F_\eta) | \psi_f(+F_\eta) \rangle}{2F_\eta} \quad (\eta = x, y, z),$$
(8)

where  $\tilde{\psi}_i$  and  $\tilde{\psi}_f$  are model vectors corresponding to initial and final electronic states, respectively.

Consider the FF calculation of TDMs for the  ${}^{2}S_{1/2} \leftrightarrow {}^{2}P_{1/2}^{o}$  and  ${}^{2}S_{1/2} \leftrightarrow {}^{2}P_{3/2}^{o}$  transitions in the Rb atom. The basis set and effective core potential employed were taken from [34]; basis set was reduced to [7s7p5d3f2g]. The simplest active space required to describe  ${}^{2}S_{1/2}$ ,  ${}^{2}P_{1/2}^{o}$  and  ${}^{2}P_{3/2}^{o}$  states of Rb comprised 5s- and 5p-spinors (overall 8 spinors, or 4 Kramers pairs).

Firstly, we perform the Hartree-Fock calculation and subsequent transformation of integrals. Note that transformed integrals over one-electron property operators (in this case, the dipole moment operator) are stored in the MDPROP unformatted file which should be copied from DIRAC scratch to the working directory:

#### pam --noarch --inp=TRA.inp --mol=Rb.mol --get="MRCONEE MDCINT MDPROP"

Other integrals are saved to the MRCONEE (one-electron Fock) and MDCINT (two-electron Coulomb) unformatted files. Note that the .PRPTRA option as well as the **\*\*PROPERTIES/.DIPOLE** specification is required in the TRA.inp input file:

```
**MOLTRA
.PRPTRA
.SCHEME
4
.ACTIVE
energy -100.0 20.0 0.5
**PROPERTIES
.DIPOLE
```

These lines are used to ask DIRAC to calculate and transform dipole moment integrals (these integrals are used later to construct the perturbation due to the external field). Also note that the external uniform electric field lowers spherical symmetry of the atom to  $C_{\infty v}$ .

Now we are ready to run FSCC calculations in EXP-T. At this step we use the oneprop directive to add an external field F contribution to the Fock operator:

$$f_{pq}' = f_{pq} + (d_x)_{pq}F_x + (d_y)_{pq}F_y + (d_z)_{pq}F_z$$

Here we use only the  $d_z$  component, stored in the MDPROP file in the ZDIPLEN record. The operator which is added to the Fock matrix must be fully symmetric. The first EXP-T input file (for field strength -F):

```
# title for the task (of type "string")
title "Rb atom -- relativistic EA-CCSD calculation"
# max number of iterations
maxiter 50
# convergence threshold (by cluster amplitudes)
conv 1e-9
# target Fock space sector
sector 0h1p
# active space specification: 8 lowest virtual spinors (5s,5p1/2,5p3/2)
nactp 8
# add the external field contribution to the Fock matrix:
# Fock = Fock + d_z * F_z
# (F_z = 1e-5)
oneprop -1e-5 0.0 "ZDIPLEN"
```

For the other field strength +F it is convenient to reuse two-electron integrals, which are already sorted (the **reuse** option). One-electron integrals depend on the finite perturbation parameter and must be recomputed. Cluster amplitudes converged at -F can also serve as a good initial guess for the calculation at +F. This will lead to significant computational savings. Use the **--no-clean** option to prevent the deletion of files containing sorted two-electron integrals and converged amplitudes.

The second EXP-T input file (for field strength +F):

```
title "Rb atom -- relativistic EA-CCSD calculation"
maxiter 50
conv 1e-9
sector 0h1p
nactp 8
```

```
reuse 2-integrals amplitudes
oneprop +1e-5 0.0 "ZDIPLEN"
```

Run EXP-T and save formatted files HEFF with effective Hamiltonian matrices to the working directory:

```
expt.x --no-clean input_F- | tee ccsd_F-.out
mv scratch/HEFF HEFF1
expt.x --no-clean input_F+ | tee ccsd_F+.out
mv scratch/HEFF HEFF2
```

The electronic spectrum is given at the end of the EXP-T output file. Corresponding lines begin with the @ symbol:

nmetry
2+ 1/2-
2+ 1/2-
2+ 1/2-
2+ 3/2-

(compare with experimental values 12578.950 cm<sup>-1</sup> and 12816.545 cm<sup>-1</sup> [32]; note that the  ${}^{2}P_{3/2}^{o}$  state is split in the external electric field).

Finally, we turn to the calculation of transition moments. The input file for the heffman.x utility is as follows (see Sec. 6.1 for details and explanation):

file: HEFF1
file: HEFF2
sector: 0h1p
rep: 1
step: 0.00002
print: 1

Run heffman.x:

heffman.x < ff.inp | tee ff.out

Transition moments and Einstein coefficients:

tran	siti	on	energy, cm^-1	•	•	d ^2	d	osc str	A,10^6 s-1
1	->	2	12555.138		•	2.627905	1.62108158	0.100220	10.537579
1	->	3	12786.559			5.173064	2.27443710	0.200921	21.911672
2	->	3	231.421		•	0.000000	0.00001167	0.00000	0.000000

Note: this is only the  $|d_z|$  component of the full transition moment d. It should be mentioned that the small, but non-zero value of |d| in the last line is an artifact of the FF approach.

Taking into account the degeneracy of the  ${}^{2}P_{3/2}^{o}$  state and the spherical symmetry of the Rb atom, the full transition moments |d| and full Einstein coefficients can be calculated:

$$A({}^{2}S_{1/2} \leftarrow {}^{2}P_{1/2}^{o}) = 31.6 \text{ s}^{-1}$$
$$A({}^{2}S_{1/2} \leftarrow {}^{2}P_{3/2}^{o}) = 32.9 \text{ s}^{-1}$$

(compare with experimental values  $34.0 \text{ s}^{-1}$  and  $37.0 \text{ s}^{-1}$  [32]).

Results can be improved significantly by extending the basis set and active space [34].

#### 5.6 Direct approach to transition dipole moment calculations: the Rb atom

For input/output files see examples/Rb\_atom\_TDMs\_direct.

The direct approach to transition moment calculations consists in the substitution of the CC wave functions  $|\psi_i\rangle = \{e^T\} |\tilde{\psi}_i\rangle$  and  $|\psi_f\rangle = \{e^T\} |\tilde{\psi}_f\rangle$  into the expression for the matrix element  $O_{if} = \langle \psi_i | O | \psi_f \rangle$ . Resulting infinite summations are truncated at some power of T. Detailed description of the theoretical background is given in [4]. Actually the same approach was previously developed for atomic applications, see the papers [35, 36, 37] (and references therein).

Let us apply this direct techique to calculate transition dipole moments in Rb. Since property matrix elements are evaluated analytically, there is no need to reduce the symmetry from  $D_{\infty h}$  to  $C_{\infty v}$ . We again begin with the DIRAC step to obtain unformatted files with transformed molecular integrals:

#### pam --inp=TRA --mol=Rb --noarch --get="MRCONEE MDCINT MDPROP"

The MDPROP file contains three entries XDIPLEN, YDIPLEN and ZDIPLEN representing matrices of the  $d_x$ ,  $d_y$  and  $d_z$  components of the dipole moment operator. For some reasons the XDIPLEN and ZDIPLEN matrix elements are purely imaginary (this was found in the series of preliminary numerical experiments), appealing to complex arithmetics in CC calculation. Such a situation is quite typical.

Thus we recommend to perform CC calculation of properties in two steps. At the first step we run the usual FS-CC calculation with the real numbers employed. EXP-T input file looks like:

```
title "Rb atom -- relativistic EA-CCSD calculation"
maxiter 50
conv 1e-9
sector 0h1p
nactp 8
disk_usage 0
```

At the second step we calculate transition moments. We should turn on complex arithmetics and restart the calculation from the pre-sorted integrals and converged amplitudes (they will be automatically expanded to the complex arithmetics):

```
title "Rb atom -- relativistic EA-CCSD calculation"
maxiter 50
conv 1e-9
sector 0h1p
nactp 8
disk_usage 0
reuse 2-integrals amplitudes
arith complex
print "model vectors"
mdprop "XDIPLEN" sym 1u+ approx 2 2 scheme hermitian
mdprop "XDIPLEN" sym 1u- approx 2 2 scheme hermitian
mdprop "YDIPLEN" sym 1u+ approx 2 2 scheme hermitian
mdprop "YDIPLEN" sym 1u- approx 2 2 scheme hermitian
mdprop "YDIPLEN" sym 1u- approx 2 2 scheme hermitian
mdprop "ZDIPLEN" sym 1u- approx 2 2 scheme hermitian
mdprop "ZDIPLEN" sym 0u approx 2 2 scheme hermitian
```

Note that the instruction print "model vectors" was added. It prints additional information about electronic energy levels and is mandatory for further successful processing of the output file with the expt\_spectrum.py script (see Sect. 6.4).

Running EXP-T:

expt.x --no-clean ccsd.inp > ccsd.out
expt.x --no-clean tdm.inp > tdm.out

The second run of EXP-T typically produces huge output files since all possible pairs of transition matrix elements are calculated. The required TDM values can be found in it manually or extracted using the expt\_spectrum.py script:

```
expt_spectrum.py ccsd.out -irep 1/2g+
```

where the -irep option defines the irrep to which the initial states belong. The output will be as follows:

```
['1/2g+']
initial irreps:
initial states:
                  all
final irreps
              :
                  all
final states :
                  all
transitions 1/2g+ -> 1/2u+
                e_i, cm^{-1}
                             e_f, cm^{-1}
                                           . . .
                                                    |d|,a.u.
                                                                 osc.str.
                            12555.1431
                                                    1.757863
   1 ->
            1
                   0.0000
                                           . . .
                                                                 0.117846
   1 ->
            2
                   0.0000
                            12786.5663
                                                    2.480855
                                                                 0.239046
                                           . . .
transitions 1/2g+ -> 1/2u-
                e_i,cm^-1
                             e_f, cm^{-1}
                                                    |d|,a.u.
                                                                 osc.str.
   1 ->
            1
                   0.0000
                            12555.1431
                                                    2.485994
                                                                 0.235693
                                           . . .
   1 ->
            2
                   0.0000
                            12786.5663
                                                    1.754230
                                                                 0.119523
                                            . .
transitions 1/2g+ -> 3/2u+
                e_i,cm^-1
                             e_f, cm^{-1}
                                                     |d|,a.u.
                                                                 osc.str.
   1 ->
            1
                   0.0000
                            12786.5663
                                                    3.038415
                                                                 0.358569
                                           . . .
```

One can see that the direct technique in the quadratic approximation (approx 2 2) overestimates transition moments by ~ 9% (compared to the finite-field approach). It can be useful to recalculate transition moments in the linear approximation (approx 1 0); for the particular case of the Rb atom it gives nearly the same answers. This is not a surprising result since cluster amplutudes are not large (< 0.07), ensuring even the linear approximation to work fine.

## 5.7 «High-spin» CCSD for open shell systems

For input/output files see examples/02\_highspin.

The example considered in this subsection is borrowed from the DIRAC program manual [17]. Let us calculate the energy of the ground (triplet) state of the  $O_2$  molecule.

The DIRAC input file contains directives for the HF calculation and subsequent integrals transformation:

(O2.mol)

```
0
       0.0
                  0.0
                           0.60376
0
       0.0
                  0.0
                           -0.60376
LARGE BASIS cc-pVDZ
FINISH
   (TRA.inp)
i
! calculation of the triplet ground state of molecular oxygen
! (Dirac-Coulomb Hamiltonian)
**DIRAC
.TITLE
molecular oxygen -- ground state
.WAVE F
.4INDEX
**WAVE FUNCTIONS
.SCF
*SCF
.CLOSED SHELL
 68
.OPEN SHELL
1
2/4,0
*END OF
```

DIRAC writes spinor occupation numbers to the MRCONEE unformatted file, but only spinors belonging to closed shell (the .CLOSED\_SHELL keyword) are labeled as occupied. This is why EXP-T will calculate the energy for the  $O_2^{2+}$  cation without manually specified occupation numbers.

EXP-T provides two ways to set the occupation numbers of spinors:

• by setting the number of occupied spinors in each irrep (the occ\_irreps keyword). For the example under consideration:

```
occ_irreps [1/2g+]:3 [1/2g-]:2 [3/2g+]:0 [3/2g-]:1
occ_irreps [1/2u+]:2 [1/2u-]:2 [3/2u+]:1 [3/2u-]:1
```

Note that the 1*s*-orbitals of oxygen atoms are frozen here, and the corresponding electrons should not be accounted for in the occ\_irreps instructions. For the example discussed:

- 3 electrons occupy the lowest spinors belonging to the 1/2g+ irrep;

- 1/2g-, 1/2u+, 1/2u-: 2 electrons in each irrep;
- 3/2g-, 3/2u+, 3/2u-: 1 electron in each irrep.
- you can also specify the occupation number for *each* spinor explicitly:

To identify spinors which should be populated in the electronic state under consideration, we again recommend starting the calculation with the default occupation numbers, then interrupting the task, finding information about spinors and restarting the calculation with the correct occupation numbers.

We recommend to use the dirac\_mointegral\_export program from DIRAC to process the MRCONEE unformatted file and get such a string of occupation numbers.

Note that, in fact, non-canonical orbitals are used, therefore the Fock matrix is not diagonal and the vacuum determinant energy  $\langle \Phi_0 | H | \Phi_0 \rangle$  is not equal to the SCF energy value obtained before:

```
-149.686661451842
Total SCF energy =
. . .
sorting one-electron integrals ...
 Fock matrix reconstruction ...
 SCF energy (energy of reference determinant) = -149.718144633810 a.u.
 SCF energy (energy of reference determinant) was updated:
   old energy =
                  -149.686661451842 a.u.
   new energy =
                   -149.718144633810 a.u.
. . .
         SCF reference energy =
                                  -149.718144633810
     CCSD correlation energy =
                                     -0.366958682652
            Total CCSD energy =
                                   -150.085103316462
```

# 6 Utility programs

## 6.1 heffman.x – manipulations with effective Hamiltonian matrices

The heffman.x program (written by A. Zaitsevskii) is designed to perform manipulations with FS-RCC effective Hamiltonians matrices. Implemented features are:

- the finite-field technique for transition dipole moments calculations [22, 34, 3];
- extraction of spin-orbit interactions from full-relativistic models by projection [5];
- Pade extrapolation of series of effective Hamiltonians [7].

heffman.x reads matrices of effective Hamiltonians from formatted files supported by both EXP-T and the latest versions of DIRAC.

- to compute finite-field transition moments [34, 3] two files with effective Hamiltonians must be provided, calculated at field 1 and field 2;
- to separate spin-orbit interactions by projection, two files with effective Hamiltonians must be provided, the first for spin-orbit (nearly) switched off and the second for spin-orbit switched on.

In this case you cannot freeze spinors at the FS-CC step!

• to perform Pade extrapolation of the series of effective Hamiltonians, you have to put a sequence of effective Hamiltonians to a single file using the cat command:

#### cat heff\_n\_eq\_1 heff\_n\_eq\_2 heff\_n\_eq\_3 > some\_file\_name

Recall that each Hamiltonian is calculated at its own attenuation parameter n. You can also perform extrapolation before TDM or SO calculation, just use such "composite" files with effective Hamiltonians.

To run heffman.x program, input file containing the list of instructions must be redirected to stdin:

heffman.x < input\_file</pre>

heffman.x input files are very similar to EXP-T input files, but keywords are separated from their arguments with the ':' symbol, and the symbol '#' at the beginning of a comment must be placed in the first position of the line.

List of instructions:

file: file with the first effective Hamiltonian or with a sequence of heff's to be extrapolated, e. g.

- heff for the first field value;
- heff for the spin-orbit interaction (nearly) switched off;
- simply heff you want to get the eigenstates;
- sequence of heff's you want to Pade-extrapolate.

#### Syntax: file: <string path-to-first-heff>

If you are doing finite-field transition moment calculations or SO separation, you have to supply the second file (see the next item).

file: heff for another field value, or heff for the spin-orbit switched on, respectively. If there is no second file, the effective Hamiltonian eigenvectors and eigenvalues are written to the formatted file HEFFEVF.

Syntax: file: <string path-to-second-heff>

**sector**: the Fock space sector you are interested in (heff files for higher sectors contain several heff matrices).

Syntax: sector: <H>h<P>p

Example: 0h1p, 1h0p, 2h0p, 0h2p, 1h1p.

rep: irreducible representation you are interested in. As a rule, rep: 1 stands for the fully symmetrical rep.

Syntax: rep: <integer n>

Enumeration of irreps begins from 1!

main: for the FF TDM calculations or Pade extrapolation: simply the number of lowest states or the range of states of given sector and symmetry you want to see in the output file. For the SO extraction by the projected approximant construction [7] this keyword defines the subspace of states used for projection.

Syntax: main: (<integer number> | <integer number1>-<integer number2>).

Default: all states.

Example: main: 10 (or, equivalently, main: 1-10)

[L/M] *(optional)* to perform Pade extrapolation, you have to indicate which Pade approximant [L/M] should be constructed ([0/1], [1/1], [0/2] etc). You should have L+M+2 effective Hamiltonian matrices to build the [L/M] approximant. You can cut off all states above subspace done by the main keyword (see above); to do this, add the cut keyword.

Syntax: [L/M] [cut]

Default: Padé extrapolation disabled

Example: [0/1] cut

step: (only for FF calculations.) Difference of field strenghts (a.u.) in two finite-field calculations. The presence of this line activates the FF TDM calculation mode (do not use this keyword in other types of calculations).

Syntax: step: <real value>

scale: (only for SO calculations.) Fraction of effective spin-orbit retained in "scalar" calculations to get good relativistic symmetries (typically of order  $10^{-4}$ ). Omit this line if you do not want to diabatize anything.

Syntax: scale: <real value>

ground: ground state energy (a.u.). Useful only when you have to compute energies with respect to some state (normally ground) which belongs to another rep or sector.

Syntax: ground: <real energy> Default: ground: 0.0 print: print level. print: 0 or print: 1 are normally ok; enlarging this number, you can get any amount of superfluous information listed.

> Syntax: print: <integer level> Default: print: 1

Bloch activates the computation of finite-field transition dipoles with Bloch effective Hamiltonian eigenvectors (note: TDM matrices will be non-Hermitian, this has no physical meaning). Otherwise, FF computations are performed with des Cloizeaux effective Hamiltonian eigenvectors ("hermitization").

Syntax: Bloch

Default: "hermitization" is enabled in finite-field calculations

When using any of the techniques implemented in the heffman.x program, please, cite the corresponding paper.

# 6.2 expt\_transform\_natorb.x - transformation of quasi-natural spinors to the AO basis

Having model vectors  $|\tilde{\psi}_i\rangle$  we can construct approximate density  $d_{pq} = \langle \tilde{\psi}_i | a_p^{\dagger} a_q | \tilde{\psi}_i \rangle$  and transition density matrices  $d_{pq} = \langle \tilde{\psi}_i | a_p^{\dagger} a_q | \tilde{\psi}_j \rangle$  and then calculate "quasi-natural" (transition) molecular spinors (NS and NTS, respectively). These operations can be performed by the EXP-T itself (see Sec. 4, keyword natorb). EXP-T stores NS and NTS expanded in the basis of molecular spinors in formatted files named NATORB\*.dat. To use visualization tools one have to transform NS or NTS to the AO basis. This transformation can be performed by the expt\_transform\_natorb.x utility.

expt\_transform\_natorb.x requires two command-line arguments:

- path to the DIRAC output file containing molecular spinors expanded in the basis of atomic orbitals;
- path to the NATORB\*.dat formatted file produced by EXP-T and stored in its scratch directory (contains natural spinors expanded in the basis of molecular spinors).

To ask DIRAC to print molecular spinors (vectors) to the output, enable the .ANALYZE module in the **\*\*DIRAC** section and then add the following text to the **\*\*ANALYZE** section of the DIRAC .inp input file [38]:

\*\*ANALYZE .PRIVEC \*PRIVEC .AOLAB .VECPRI all

(.AOLAB = "print vectors in the AO basis"). Note that for point groups with inversion one must add the second string specifying the range of molecular spinors (e. g. all).

Running expt\_transform\_natorb.x (for example):

expt\_transform\_natorb.x dirac.out NATORB\_0h1p\_1:1.dat > nts.out

The transformed spinors are flushed to **stdout** again in the DIRAC four-column format and can be redirected to file.

## 6.3 expt2pam.x – input file format converter

EXP-T does not provide subroutines for solution of Hartree-Fock equations and integral transformation, it completely relies on molecular integrals calculated with the help of any other quantum chemistry software. Thus EXP-T input file do not contain any information on molecular geometry, basis set and pseudopotentials. However, to simplify interaction with the third-party software, the special file format was developed. It is designed to be quite flexible and free of any format restrictions typical for software written in Fortran.

The EXP-T-style input files containing information on molecular geometry, basis sets and pseudopotentials are recognized by the expt2pam.x program, which can produce input files for the following software:

- DIRAC [1, 2]: option -dirac;
- CFOUR [39]: option -cfour;
- OneProp [6]: option -oneprop.

The input file format closely resembles that of NWChem [40, 41, 42, 43], but is substantially simplified. The most important points on the format to be emphasized are:

- only Cartesian coordinates for molecular geometry definition are available to the moment;
- primitive Gaussian functions can be defined without the superfluous contraction coefficient 1.0;
- primitive Gaussian functions representing the uncontracted part of the basis set with the given angular momentum L can be placed into the same L-block.

Below the example of such an input file demonstrating the main features of the expt2pam format is presented (the UO<sub>2</sub> molecule [44]):

```
#
# Uranium dioxide molecule UO2
#
geometry units angstrom
    0.0 0.0
                0.000
  U
    0.0 0.0
                1.775
  0
  0
     0.0 0.0
               -1.775
end
#
 Contracted basis set adapted for the 28-in-core GRPP by N.S. Mosyagin
#
#
basis
US
 11931.71300000
  6539.04200000
  3678.23270000
  2139.54940000
  1288.45330000
   792.07718000
   484.69932000
   298.00127000
   186.73959000
   120.33238000
```

## 75.85031400 49.22312100 30.97937000 19.33898600 11.90917900 7.28108690 4.52381560 2.69147870 1.62933420 1.02810670 0.60588439 0.35564698 0.20180019 0.12481356

0.06383616

UΡ

0 F

8256.41640000 4196.77840000 2191.89420000 1174.82830000 644.61478000 361.41388000 206.27034000 119.33670000 70.14308600 41.70407000 24.70288000 14.49639800 8.57587000 4.98086080 2.84040830 1.59903720 0.86476597 0.45724205 0.23893352 0.11879300 0.05700662

U D

6012.79360000 2754.36670000 1351.96790000 699.79805000 377.01141000 209.73941000 119.30191000 69.11171400 40.43241500 23.65423900 13.806399008.06231670

	4.66375400					
	2.63565040					
	1.45796000					
	0.77481216					
	0.37248472					
	0.17195534					
	0.07613361					
UF						
11	63.36020000					
5	573.50499000					
2	285.71163000					
1	36.47329000					
	69.43427500					
	36.84465700					
	19.84804900					
	10.72168500					
	5.70450120					
	2.92433440					
	1.43960400					
	0.67385930					
	0.29434970					
	0.11323985					
UG						
	28.78919130	0.00068974	-0.00118120	-0.00737416	0.00494647	-0.10867784
	14.01320380	-0.00898795	-0.05421964	0.17284128	-0.36594951	-0.33163329
	6.82095857	0.07975524	0.18134441	-0.50920321	1.20716202	1.79274020
	3.32011698	0.29993353	0.65729862	-0.50617914	-0.67702022	-2.65141994
	1.61607443	0.39302861	0.04007231	1.01049051	-0.71780643	2.66362756
	0.78662787	0.34151054	-0.44799971	0.05077223	1.59628888	-2.01035588
	0.38289289	0.18576214	-0.38383279	-0.74551549	-1.12509753	0.90492997
UΗ						
	14.01320380	-0.02449353	-0.07337288	-0.15906456	-0.23132505	
	6.82095857	0.10594999	0.30442614	0.71623431	1.22980099	
	3.32011698	0.35765256	0.65468802	0.09614021	-1.24473080	
	1.61607443	0.43019757	-0.18764249	-0.95514471	0.32540556	
	0.78662787	0.31957043	-0.48353526	0.41946211	0.73545263	
	0.38289289	0.10197395	-0.22548628	0.48672026	-0.90679010	
υI						
	6.82095857	0.12524957	-0.37984959	0.81416792		
	3.32011698	0.42961205	-0.61109425	-0.35518331		
	1.61607443	0.42313291	0.38916709	-0.73501069		
	0.78662787	0.31396902	0.50454838	0.88347742		
ond						

end

#

```
# Relativistic pseudopotential (RPP) for Uranium by N.S. Mosyagin
# 28e in core, QED+Breit are included
#
ecp
```

U nelec 28

U UL

1	573.3050904771354	-11.40084159918683
1	192.8309334143903	-7.097553073552547
1	59.69944003435170	-28.01857233638748
1	41.19441327874497	42.95299466064716
1	27.48466038227223	-39.29896548592660
1	17,41575927233871	21.07523831226354
1	10.74156012298913	-8.338244516015038
1	6 516110723607048	3 069168416643151
1	4 072460983953957	-1 269035430883378
1	2 593752889495458	0 4038040550003643
1	1 642108539873937	_0 7510043731679161F_001
1	0 5908313234915240	0 3890633153407208F_002
יד	0.0000010204010240	0.30300331334072001-002
0.0	107 0208350128211	6 000000000000000
1	120 5661840070536	355 1867566615011
1 1	03 02818620674168	
1 1	66 12102012007649	1657 092006766264
1	45 26004796250122	
1	40.000000000000000000000000000000000000	1304.220230300319
1	31.03973237770222	-440.9309445470000
1	21.1002000420004	7 27052000267057
1	8.842323178530615	-7.379538028307257
T	2.503929576228269	3.473180786461057
	2.925942781684598	-5.455412335112918
UP	0.011400755000100	F 0000000000000000
0	8.311409755029183	5.00000000000000
1	126.52/431986/533	-10.06000092823524
1	85.18677663478607	70.55297376279486
1	51.21216731114048	-221.80/3/35/31159
1	29.49108937084788	312.1302683559188
1	17.84600190857478	-299.1627976468299
1	5.463491362893393	229.1321036359295
2	8.266178701884957	-440.5446582608856
2	5.564984596440512	-230.2801015020814
2	3.621496981299310	0.6144684865466843
υD		
0	63.75445017844640	3.000000000000000
1	35.83829800943973	-244.0643648227415
1	25.77961295248244	369.6289272300128
1	16.62718251626653	-66.05286509307552
2	21.10083097720674	-511.0304327796694
2	13.00689068383810	209.7715531529684
2	7.652250767129301	-56.31696437943174
2	4.429977591227489	19.86068552036704
2	2.777850799855992	-8.087414881056247
2	1.804331049757116	2.539779591468687
2	1.174834421468355	-0.5000680379316691
2	0.7571835654831253	0.4510869608679968E-001

 $\operatorname{end}$ 

# # spin-orbit part of RPP for Uranium # so UΡ 126.5274319867533 -60.8272897577457000 1 1 85.18677663478607 144.881442506022000 1 51.21216731114048 -174.959781119600000 1 29.49108937084788 113.392635098503000 17.84600190857478 -60.8426035265438000 1 1 5.463491362893393 45.8248263173737000 2 8.266178701884957 -82.3448048559949000 2 5.564984596440512 -47.9408659655601000 2 3.621496981299310 0.4919304400803720 UD 1 35.83829800943973 -0.4757330698863370 25.77961295248244 -8.8286299074417900 1 1 16.62718251626653 9.6090737588470800 2 21.10083097720674 14.829424758465000 2 13.00689068383810 -30.085848351789200 2 7.652250767129301 10.782019233677400 2 4.429977591227489 -3.9646231214668700 2 2.777850799855992 1.4116480556389400 2 1.804331049757116 -0.3371713502943750 2 1.174834421468355 0.0452391741906467 2 0.7571835654831253 -0.0020991859739297 UΕ 573.3050904771354 1.9478068936173300 1 1 192.8309334143903 0.3189229591513530 1 59.69944003435170 1.1230953284521100 41.19441327874497 1 -2.2059715944841200 1 27.48466038227223 2.4430555431733500 17.41575927233871 -1.79281906937415001 1 10.74156012298913 1.0519241531739200 1 6.516110723607048 -0.5026634876599670 1 4.072460983953957 0.2165908956875290 2.593752889495458 1 -0.0639500309090571 1 1.642108539873937 0.0107075321136888 1 0.5908313234915240 0.0000807064226547 end

#

```
"
# basis for oxygen: aug-cc-pVQZ-DK
#
basis
0 S
61420.00000000 0.00019890 -0.00004520
```

	919	99.0000000	0.0008884	- 04	0.00020	260
	209	91.00000000	0.0039425	50 -	0.00089	480
	59	90.9000000	0.0156072	20 -	0.00360	580
	19	92.30000000	0.0529461	.0 –	0.01230	330
	e	59.32000000	0.1464382	20 -	0.03646	840
		26.97000000	0.3054148	30 -	0.08316	710
	-	11.10000000	0.3979979	90 -	0.15217	400
	-	4 68200000	0 2164354	LO -	0 11462	710
		1 42800000	0 0175357	20 20	0.29002	590
		0 55470000	_0 0024923	20	0.58584	370
		0.33470000		70	0.0000	2400
0	C	0.20070000	0.0009307	0	0.21013	9420
U	a	1 4000000				
		1.42800000				
		0.55470000				
		0.20670000				
		0.06959000				
0	Р					
	6	53.42000000	0.0061107	0		
	-	14.66000000	0.0418957	0		
		4.45900000	0.1612329	90		
		1.53100000	0.3566752	20		
		0.53020000	0.4481837	0		
		0.17500000	0.2450542	20		
0	Р					
		1.53100000				
		0.53020000				
		0.17500000				
		0.05348000				
0	D					
_		3.77500000				
		1.30000000				
		0 44400000				
		0 15400000				
п	F	0.10400000				
U	1	2 66600000				
		2.00000000				
		0.85900000				
0	a	0.32400000				
U	G	1 04600000				
		1.84600000				
		0.71400000				
er	ıd					
#	_					
#	Empty-co	ore semilocal	RPP for ox	ygen	by N.S.	Mosyagin
#						
e	ср					
0	nelec O					
0	UL					
	1	2.44309772428	80e+02	1.36	9170316	080e-01

1	8.520362677383e+01	-2.852592197455e-01
2	1.054731223008e+02	2.684494987739e+00
2	1.825657247757e+01	-9.433884948610e-03
2	4.806589859822e+00	-1.366233123930e-03
2	1.753580551235e+00	4.456314085999e-04
2	6.566864682586e-01	-2.781240176237e-04
2	4.686645649504e-01	3.558106708362e-04
0 S		
1	1.376748432210e+02	5.591554374814e-02
1	1.378552088359e+01	3.868632083368e-03
2	6.768869843542e-01	5.419543370664e-04
2	4.618614187585e-01	-3.882304906071e-04
end		
SO		
ΟP		
1	2.443097724280e+02	-9.027082481378e-02
1	8.520362677383e+01	2.802454585393e-01
2	1.054731223008e+02	-2.610156015857e+00
2	1.825657247757e+01	1.266811741578e-02
2	4.806589859822e+00	2.570019532778e-03
2	1.753580551235e+00	2.973424726019e-04
2	6.566864682586e-01	6.007087622745e-05
2	4.686645649504e-01	-5.673502079255e-05
end		

# 6.4 expt\_spectrum.py – a program for extraction of transition moments from EXP-T output files

EXP-T calculations of transition matrix elements can produce huge output files. To simplify their analysis and extraction of subsets of matrix elements needed for a particular problem, the expt\_spectrum.py script is provided. It is written in Python 3 and requires the pre-installed NumPy library.

Options of the expt\_spectrum.py script are as follows:

```
$ expt_spectrum.py --help
usage: expt_spectrum.py [-h] [-u UPPER] [-irep INITIAL_IRREP] [-frep FINAL_IRREP]
    [-i INITIAL_STATE INITIAL_STATE] [-f FINAL_STATE FINAL_STATE]
    [-H] [-s] [-ev] filename
```

parses exp-t output files and collects information about electronic transitions and their dipole moments

```
positional arguments:
    filename
```

optional arguments:	
-h,help	show this help message and exit
-u UPPER	upper energy limit for the electronic states considered (cm <sup>2</sup> -1 or ev)
-irep INITIAL_IRREP	irrep of initial electronic states

-frep FINAL_IRREP	irrep of final electronic states
-i INITIAL_STATE INIT	IAL_STATE
	initial electronic state
-f FINAL_STATE FINAL_S	STATE
	final electronic state
-H	force hermiticity of transition moments
-S	print detailed information on electronic spectrum
-ev	use electron-volts units for energy

Names of irreducible representations are common for the EXP-T package and are listed in Appendix C.

For example, let us consider allowed E1-transitions in the AcF molecule. Its electronic states were thoroughly studied in [19]. Suppose that calculated transition moments are stored in the EXP-T output file named tdms.out. The following command should be used to show all possible allowed transitions from the ground state ( $\Omega = 0^+$ ):

#### expt\_spectrum.py tdms.out -i 0 1

Note that DIRAC does not discern  $0^+$  and  $0^-$  states, they both are denoted as 0. The ground state is the first (1) in the irrep with name 0.

To calculate radiation lifetime of an excited state and estimate branching ratios during its decay, one should obtain information on transition dipole moments between this excited state and all states lying below it. For the particular case of the AcF molecule, the most strong transition from the  $X0^+$  ground state is the transition to the (8)1 state. Thus this excited state seems to be the most accessible in experiment. To extract all information about transitions from (8)1 to the low-lying states, one can use the following command:

#### python expt\_spectrum.py tdms.out -i 1+ 8 -u 26200

Here the -u option is used to restrict the range of final states to those lying below (8)1 ( $\approx 26170 \text{ cm}^{-1}$ ).

The expt\_spectrum.py routine can be easily combined with Unix grep and awk utilities to easily extract transition dipole moment functions or even surfaces from a large set of EXP-T output files.

# 6.5 expt\_diatomic.x – rotational-vibrational levels and properties of diatomic molecules

If one has a potential energy curve U(R) for a diatomic molecule, one can solve the radial Schrödinger equation and obtain rovibrational levels and wavefunctions as well as any other properties (molecular constants, transition moments, Franck-Condon factors). The radial Schrödinger equation is written as:

$$\left[-\frac{\hbar^2}{2\mu}\frac{d^2}{dR^2} + \frac{J(J+1)}{2\mu R^2} + U(R)\right]\psi_{vJ}(R) = E_{vJ}\ \psi_{vJ}(R),\tag{9}$$

where

•  $\mu = \frac{M_1 M_2}{M_1 + M_2}$  – reduced mass,

- v vibrational quantum number,
- J rotational quantum number.

Normally this second-order differential equation is solved numerically using some mesh of radial points (see, for example, [45, 46, 47]). Within the current implementation, the U(R) function is interpolated using the cubic natural spline. The matrix version of either the Numerov method [48] or the 2nd order finite-difference method [45] is used to perform integration. Only bound states can be treated to the moment.

#### 6.5.1 Input file syntax

The syntax of input files for the solver is quite similar to that of the coupled-cluster program. Single line comments start with the '#' symbol. The following directives are allowed:

masses masses of isotopes  $M_1$  and  $M_2$  (in amu).

Syntax: masses <float M1> <float M2>

charge net charge of the molecule q. For charged molecules the "charge-modified reduced mass" formula [49, 47] is used:

$$\mu = \frac{M_1 M_2}{M_1 + M_2 - m_e q},\tag{10}$$

where  $m_e$  stands for the electron mass. Can be useful for light molecular ions.

Syntax: charge <integer q>

- viblevels range  $[v_{min}, v_{max}]$  of vibrational levels of interest. Syntax: viblevels <integer vmin> <integer vmax> Default: viblevels 0 0
- rotlevels range  $[J_{min}, J_{max}]$  of rotational levels of interest. Syntax: rotlevels <integer Jmin> <integer Jmax> Default: rotlevels 0 0
- grid\_size number of radial mesh points.

Syntax: grid\_size <integer npoints>

Default: grid\_size 300

solver integration algorithm:

solver numerovNumerov method [48, 47] (default)solver fd22nd order finite-difference method [45]

mapping meshkov08 <float ra in Angstroms> <float beta>

mapping formula for the radial variable (disabled by default). Reduced variable mapping first proposed by V. V. Meshkov and co-authors [45]:

$$y(r) = \frac{r^{\beta} - r_{a}^{\beta}}{r^{\beta} + r_{a}^{\beta}} \quad y \in [-1, 1]$$
(11)

Mapping is available only for the finite-difference solver (solver fd2).

- write\_psi flush wavefunctions  $\psi_{Jv}(R)$  to the formatted text files (disabled by default).
- potential point-wise potential energy curves and/or property matrix element. The first (r) and the second  $(U_1(r))$  columns are mandatory. The third (property) and the fourth  $(U_2(r))$  columns are optional. If two potentials are specified, transition matrix elements will be calculated. To calculate Franck-Condon factors, use M(r) = 1 in the third column. Syntax:

potential <units of distance> <units of energy>
 <float r> <float U1(r)> [<float M(r)> [<float U2(r)>] ]
 . . .
end

Allowed units of distance are: angstrom, atomic, bohr, au.

Allowed units of energy are: cm-1, cm, atomic, hartree.

#### 6.5.2 Example: Franck-Condon factors for the AcOH<sup>+</sup> molecular ion

To estimate the suitability of the molecule for the direct laser cooling one should calculate vibrational branching ratios and Franck-Condon factors for the working electronic transition. Recently the AcOH<sup>+</sup> molecular ion was proposed as a candidate for direct laser cooling [50].

This molecule is linear in both the ground (1)1/2 and the first excited (1)3/2 electronic states. If we neglect the anharmonic couplings between stretching and bending modes, the total FCF between the  $(v'_1, v'_2, v'_3)$  and the  $(v''_1, v''_2, v''_3)$  vibrational states can be factorized:

$$FCF = |\langle v_1' v_3' | v_1'' v_3' \rangle|^2 \times |\langle v_2' | v_2'' \rangle|^2$$

$$\tag{12}$$

The latter factor is the one-dimensional bending mode FCF and it is expected to be very close to unity since the corresponding sections of potential energy surfaces are perfectly parallel to each other. The overlap intergal over the  $v_3$  stretch mode is expected to be nearly equal to unity and, which is more important, this factor will be the same for the  $(v'_1, v'_2, 0)$  and  $(v''_1, v''_2, 0)$  vibrational states analyzed here since this stretching mode is not excited in the states under consideration. Thus the total FCF can be approximated with the squared one-dimensional  $\langle v'_1 | v''_1 \rangle$  overlap integral. To calculate this integral one should solve the radial Schrödinger equation, considering the OH group as a pseudoatom.

The input file for the expt\_diatomic.x program:

```
#
# calculation of Frank-Condon factors for the (1)1/2-(1)3/2 transition in AcOH+
#
 potential curves were calculated by the relativistic FS-CCSD(0h1p) method,
#
# for more details, see:
# A. Oleynichenko, L. V. Skripnikov, A. V. Zaitsevskii, V. V. Flambaum
# Laser-coolable AcOH+ ion for CP-violation searches (2022)
#
 arXiv:2112.02307 [physics.atom-ph]
#
#
           Ac
                      OH
masses 227.028000 17.003000
rotlevels 0 0
viblevels 0 5
# r
      potential1(r)
                       property(r)=1.0
                                          potential2(r)
# units: dist
                 energy
potential atomic atomic
0.340150702E+01
                -0.451045595E+03
                                    1.0 -0.451004746E+03
0.349599333E+01 -0.451064998E+03
                                    1.0 -0.451025470E+03
0.359047964E+01 -0.451079018E+03
                                    1.0 -0.451040834E+03
0.368496594E+01 -0.451088543E+03
                                    1.0 -0.451051696E+03
                                    1.0
0.377945225E+01 -0.451094322E+03
                                        -0.451058791E+03
0.387393856E+01 -0.451096994E+03
                                    1.0
                                        -0.451062746E+03
```

0.396842486E+01	-0.451097103E+03	1.0	-0.451064095E+03
0.406291117E+01	-0.451095111E+03	1.0	-0.451063297E+03
0.415739747E+01	-0.451091413E+03	1.0	-0.451060740E+03
0.425188378E+01	-0.451086341E+03	1.0	-0.451056758E+03
0.434637009E+01	-0.451080177E+03	1.0	-0.451051631E+03
0.444085639E+01	-0.451073161E+03	1.0	-0.451045598E+03
0.453534270E+01	-0.451065491E+03	1.0	-0.451038858E+03
end			

Run the job:

expt\_diatomic.x AcOH+\_FCF.inp | tee AcOH+\_FCF.out

At the end of the output file one can find the table with transition moments (and their squares). We select only the FCFs with the wavenumber of the upper state v'' = 0:

J,	J,,	v'	v''			delta e,cm^-1	prop	prop ^2
\$ 0	0	0	0		•	7279.8734	-0.942947	0.889149
\$ 0	0	1	0			6627.6841	0.325405	0.105888
\$ 0	0	2	0			5979.2731	0.069678	0.004855
\$ 0	0	3	0	•	•	5334.3122	0.010337	0.000107
\$ 0	0	4	0			4692.7001	0.001070	0.00001
\$ 0	0	5	0			4054.6587	-0.000079	0.000000

The sum of first three FCFs (see the last column) is equal to 0.9999, thus allowing one to conclude that the closed optical cycle can exist for the  $AcOH^+$  molecular cation [50].

# Appendix A Alphabetical index of keywords

```
active 15
analyt_prop 17
\texttt{arith}\ 12
compress 23
conv 20
cuda 25
damping 20
degen_thresh 12
density 17
diis 20
disk_usage 23
div_thresh 20
flush 20
hermit 17
ih_imms 15
ih_imms : frontier_energy 17
ih_imms : main_occ 16
ih_imms : npower 16
ih_imms : scale 17
ih_imms : sectors 16
ih_imms : shift_type 16
ih_imms :
            subspace 16
integrals 23
maxiter 21
mdprop 17
memory 24
model 13
mstdm 19
\texttt{nacth } 14
nactp 14
natorb 19
nroots 21
nthreads 24
\operatorname{occ} 14
occ_irreps 14
oneprop 23
openmp 25
print 12
reuse 21
roots_cutoff 21
sector 13
shift 21
shifttype 22
skip 22
spinor_labels 13
tilesize 24
title 12
txtprop 19
```

x2cmmf 23

# Appendix B Temporary files

During the calculation, EXP-T operates with temporary files stored in the *scratch* directory. We recommend to remove these temporary files after every calculation (except for those planned to be used in the future, like the HEFF files or files with cluster amplitudes). All types of temporary files generated by EXP-T are listed and described below.

- HINT one-electron integrals core Hamiltonian matrix + one-particle part of perturbation (if presented). Recommended for removal after calculation.
- VINT-\*-\*-\* two-electron integrals. These files are generated during the integral sorting stage. Recommended for removal after calculation.

Approximate number of files:  $\sim N_b^4$  ( $N_b$  – number of blocks into which the whole set of spinors is divided by symmetry and the **tilesize** parameter)

\*.sb single symmetry block of molecular integrals/cluster amplitudes. It is just a part of a diagram. During the calculation each file is mapped to some structure in RAM. Recommended for removal after calculation.

Approximate number of files:  $\sim 10 \times N_b^4$  for CCSD,  $\sim N_b^6$  for CCSDT-n and CCSDT models – can be very large!

- \*.dg file containing diagram. Diagram files containing cluster amplitudes (tlc.dg, t2c.dg etc) can be used in subsequent calculations when constructing the initial approximation to amplitudes (see Sec. 4, keyword reuse).
- HEFF formatted file containing effective Hamiltonian matrices. These files can be read by the heffman.x utility program used to perform Padé extrapolation of series of effective Hamiltonians [7], finite-field transition moments calculations [34, 3], spin-orbit coupling calculations [5].
- \*DIPLEN formatted files with dipole moment integrals (just its electronic part) in the basis of molecular spinors.
- MVCOEF\*\* binary files containing model vectors expanded in the basis of model determinants. These files are used to organize data flow inside EXP-T in the most convenient and logical manner; model vectors are used for calculation of model-space approximations of TDMs and "quasi-natural" spinors, etc.
- NATORB\*.dat formatted text files containing "quasi-natural" spinors expanded in the basis of molecular spinors. These expansions can be transformed to the AO basis with the mknatorb.x utility program (see Sec. 6.2).

# Appendix C Symmetry and irreducible representations

This section lists the names of irreducible representations (irreps) used in EXP-T.

- For nonrelativistic groups, irrep is determined by the spatial symmetry (Mulliken notation is used) and the  $M_s$  value (for  $M_s \ge 2$  only symbol 2 is used);
- For relativistic double groups we use Mulliken-type notation (completely inherited from DIRAC);
- For  $U(1) = \overline{C_{\infty v}}$  and  $U(1) \times C_i = \overline{D_{\infty h}}$  irreps are determined by  $\Omega$  values.

 $C_1$ 

 $C_2$ 

 $C_s$ 

A\_a

A\_b A\_-3/2 A\_+3/2  $A_0$  $A_4$ A\_+1 A\_-1 A\_a B\_a A\_b B\_b A\_-3/2 B\_-3/2 A\_+3/2 B\_+3/2  $A_0$ B\_0 A\_2 B\_2 A\_+1 B\_+1  $A_-1$ B\_-1 A'\_a A"\_a A'\_b A"\_b A'\_-3/2 A"\_-3/2 A'\_+3/2 A"\_+3/2 A'\_O A"\_0" A'\_2 A"\_2" A'\_+1 A"\_+1 A'\_-1 A"\_-1

 $C_i$ 

Ag_a	Au_a
Ag_b	Au_b
Ag3/2	Au3/2
Ag_+3/2	Au_+3/2
Ag_O	Au_O
Ag_2	Au_2
Ag_+1	Au_+1
Ag1	Au1

# $C_{2v}$

A1_a	B2_a	B1_a	A2_a
A1_b	B2_b	B1_b	A2_b
A13/2	$B2_{-3/2}$	$B1_{-3/2}$	A23/2
A1_+3/2	B2_+3/2	B1_+3/2	A2_+3/2
A1_0	B2_0	B1_0	A2_0
A1_2	B2_2	B1_2	A2_2
A1_+1	B2_+1	B1_+1	A2_+1
A11	B21	B11	A21

# $C_{2h}$

Bg_a	Bu_a	Au_a
Bg_b	Bu_b	Au_b
Bg3/2	Bu3/2	Au3/2
Bg_+3/2	Bu_+3/2	Au_+3/2
Bg_0	Bu_O	Au_O
Bg_2	Bu_2	Au_2
Bg_+1	Bu_+1	Au_+1
Bg1	Bu1	Au1
	Bg_a Bg_b Bg3/2 Bg_+3/2 Bg_0 Bg_2 Bg_+1 Bg1	Bg_a       Bu_a         Bg_b       Bu_b         Bg3/2       Bu3/2         Bg_+3/2       Bu_+3/2         Bg_0       Bu_0         Bg_2       Bu_2         Bg_+1       Bu_+1         Bg1       Bu1

# $D_2$

A_a	B3_a	B1_a	B2_a
A_b	B3_b	B1_b	B2_b
A3/2	B33/2	$B1_{-3/2}$	B23/2
A_+3/2	B3_+3/2	B1_+3/2	B2_+3/2
A_0	B3_0	B1_0	B2_0
A_2	B3_2	B1_2	B2_2
A_+1	B3_+1	B1_+1	B2_+1
A1	B31	B11	B21

 $D_{2h}$ 

	Ag_a Ag_b Ag3/2 Ag_+3/2 Ag_0 Ag_2 Ag_2 Ag_+1 Ag1	B1u_a B1u_b B1u3/2 B1u_+3/2 B1u_0 B1u_2 B1u_+1 B1u1	B2u_a B2u_b B2u3/2 B2u_+3/2 B2u_0 B2u_2 B2u_2 B2u_+1 B2u1	B3g_a B3g_b B3g3/2 B3g_+3/2 B3g_0 B3g_2 B3g_2 B3g_+1 B3g1	B3u_a B3u_b B3u3/2 B3u_+3/2 B3u_0 B3u_2 B3u_2 B3u_+1 B3u1	B2g_a B2g_b B2g3/2 B2g_+3/2 B2g_0 B2g_2 B2g_2 B2g_+1 B2g1	B1g_a B1g_b B1g3/2 B1g_+3/2 B1g_0 B1g_2 B1g_+1 B1g1	Au_a Au_b Au3/2 Au_+3/2 Au_0 Au_2 Au_2 Au_+1 Au1	
$\overline{C_1}$									
	A a								
$\overline{C_2}$									
	1E 2E a b								
$\overline{C_s}$									
	1E 2E a b								
$\overline{C_i}$									
	AG AU ag au								
$\overline{C_{2v}}$									
	1E 2E a b								
$\overline{D_2}$									
	1E 2E a b								
$\overline{D_{2h}}$									
	1Eg 2Eg ag bg	1Eu 2Eu au bu							

 $\overline{C_{\infty v}}$ 

1/2+	1/2-	3/2+	3/2-	5/2+	5/2-	7/2+	7/2-
9/2+	9/2-	11/2+	11/2-	13/2+	13/2-	15/2+	15/2-
17/2+	17/2-	19/2+	19/2-	21/2+	21/2-	23/2+	23/2-
25/2+	25/2-	27/2+	27/2-	29/2+	29/2-	31/2+	31/2-
0	1+	1-	2+	2-	3+	3-	4+
4-	5+	5-	6+	6-	7+	7-	8+
8-	9+	9-	10+	10-	11+	11-	12+
12-	13+	13-	14+	14-	15+	15-	16+

 $\overline{D_{\infty h}}$ 

1/2g+	1/2g-	3/2g+	3/2g-	5/2g+	5/2g-	7/2g+	7/2g-
9/2g+	9/2g-	11/2g+	11/2g-	13/2g+	13/2g-	15/2g+	15/2g-
1/2u+	1/2u-	3/2u+	3/2u-	5/2u+	5/2u-	7/2u+	7/2u-
9/2u+	9/2u-	11/2u+	11/2u-	13/2u+	13/2u-	15/2u+	15/2u-
0g	1g+	1g-	2g+	2g-	3g+	3g-	4g+
4g-	5g+	5g-	6g+	6g-	7g+	7g-	8g+
Ou	1u+	1u-	2u+	2u-	3u+	3u-	4u+
4u-	5u+	5u-	6u+	6u-	7u+	7u-	8u+

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